

Post-Newtonian N-body Codes

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Introduction

N-body tools

Three-body formulation

PN implementations

Numerical examples

Discussion

Hermite Integration

Taylor series for \mathbf{F} and $\mathbf{F}^{(1)}$

$$\mathbf{F} = \mathbf{F}_0 + \mathbf{F}_0^{(1)} t + \frac{1}{2} \mathbf{F}_0^{(2)} t^2 + \frac{1}{6} \mathbf{F}_0^{(3)} t^3$$

$$\mathbf{F}^{(1)} = \mathbf{F}_0^{(1)} + \mathbf{F}_0^{(2)} t + \frac{1}{2} \mathbf{F}_0^{(3)} t^2$$

Prediction

$$\mathbf{r}_j = \left(\left(\frac{1}{6} \mathbf{F}_0^{(1)} \delta t'_j + \frac{1}{2} \mathbf{F}_0 \right) \delta t'_j + \mathbf{v}_0 \right) \delta t'_j + \mathbf{r}_0$$

$$\mathbf{v}_j = \left(\left(\frac{1}{2} \mathbf{F}_0^{(1)} \delta t'_j + \mathbf{F}_0 \right) \delta t'_j + \mathbf{v}_0 \right); \quad \delta t'_j = t - t_0$$

New forces $\mathbf{F}, \mathbf{F}^{(1)}$

Higher derivatives

$$\mathbf{F}_0^{(3)} = (2(\mathbf{F}_0 - \mathbf{F}) + (\mathbf{F}_0^{(1)} + \mathbf{F}^{(1)}) t) \frac{6}{t^3}$$

$$\mathbf{F}_0^{(2)} = (-3(\mathbf{F}_0 - \mathbf{F}) - (2\mathbf{F}_0^{(1)} + \mathbf{F}^{(1)}) t) \frac{2}{t^2}$$

Corrector for i

$$\Delta \mathbf{r}_i = \frac{1}{24} \mathbf{F}_0^{(2)} \Delta t^4 + \frac{1}{120} \mathbf{F}_0^{(3)} \Delta t^5$$

$$\Delta \mathbf{v}_i = \frac{1}{6} \mathbf{F}_0^{(2)} \Delta t^3 + \frac{1}{24} \mathbf{F}_0^{(3)} \Delta t^4$$

AC Neighbour Scheme

Total force
$$\mathbf{F}(t) = \sum_{j=1}^n \mathbf{F}_j + \mathbf{F}_d(t)$$

Prediction

$$\mathbf{F}(t) = \mathbf{F}_n + \dot{\mathbf{F}}_d(t - t_0) + \mathbf{F}_d(t_0)$$

$$\dot{\mathbf{F}} = \dot{\mathbf{F}}_n + \dot{\mathbf{F}}_d$$

Time-scales

$$\Delta t_n \ll \Delta t_d, \quad n \ll N$$

Neighbour sphere
$$R_s^{\text{new}} = R_s^{\text{old}} \left(\frac{n_p}{n} \right)^{1/3}$$

Neighbour selection
$$|\mathbf{r}_i - \mathbf{r}_j| < R_s$$

Derivative corrections
$$\mathbf{F}_{ij}^{(2)}, \mathbf{F}_{ij}^{(3)}$$

Basic KS Relations

New coordinates $R = u_1^2 + u_2^2 + u_3^2 + u_4^2$

Time transformation $dt = R d\tau$

Coordinate transformation $\mathbf{R} = \mathcal{L}(\mathbf{u}) \mathbf{u}$

Generalized Levi–Civita matrix

$$\mathcal{L}(\mathbf{u}) = \begin{bmatrix} u_1 & -u_2 & -u_3 & u_4 \\ u_2 & u_1 & -u_4 & -u_3 \\ u_3 & u_4 & u_1 & u_2 \\ u_4 & -u_3 & u_2 & -u_1 \end{bmatrix}$$

Inverse relations

$$R_1 = u_1^2 - u_2^2 - u_3^2 + u_4^2$$

$$R_2 = 2(u_1 u_2 - u_3 u_4)$$

$$R_3 = 2(u_1 u_3 + u_2 u_4)$$

$$R_4 = 0$$

Velocity $\dot{\mathbf{R}} = \frac{2\mathcal{L}\mathbf{u}'}{R}, \quad R' = t'' = 2\mathbf{u} \cdot \mathbf{u}'$

Equations of motion

$$\mathbf{u}'' = \frac{1}{2} h \mathbf{u} + \frac{1}{2} R \mathcal{L}^T \mathbf{P}$$

$$h' = 2 \mathbf{u}' \cdot \mathcal{L}^T \mathbf{P}$$

$$t' = \mathbf{u} \cdot \mathbf{u}$$

KS Decision-Making

Close encounter	$R_{\text{cl}} = \frac{4r_{\text{h}}}{N C^{1/3}}, \quad \Delta t_{\text{cl}} = \beta \left(\frac{R_{\text{cl}}^3}{\bar{m}} \right)^{1/2}$
Time-step criterion	$\Delta t_k < \Delta t_{\text{cl}}$
Neighbour list search	list all $r_{kj}^2, \quad \Delta t_j < 2 \Delta t_{\text{cl}}$
Two-body selection	$R_{kl} < R_{\text{cl}}, \quad \dot{R}_{kl} < 0$
Dominant motion	$\frac{m_k + m_l}{R_{kl}^2} > \frac{m_k + m_j}{R_{kj}^2}$
KS initialization	$\mathbf{F}_U, \mathbf{F}'_U, \Delta\tau \ \& \ t^{(n)} \Rightarrow \Delta t$
Initialization of c.m.	$\mathbf{r}_{\text{cm}} = \frac{m_k \mathbf{r}_k + m_l \mathbf{r}_l}{m_k + m_l}$
Perturber search	$r_{\text{p}} < \left(\frac{2m_{\text{p}}}{m_{\text{b}} \gamma_{\text{min}}} \right)^{1/3} a (1 + e)$
Slow-down adjustment	$\gamma < \gamma_0, \quad \Delta\tau \Rightarrow \kappa \Delta t$
Termination test	$R > R_0, \quad \gamma > \gamma^*$
Delayed termination	$T_{\text{block}} - t > \Delta t_i$
Final iteration	$\Delta\tau$ from $\dot{\tau}, \ddot{\tau}, \dots$ and δt
Polynomial initialization	$\mathbf{F}_j, \dot{\mathbf{F}}_j, \Delta t_j, \quad j = k, l$

Hermite KS with block-steps

Predict \mathbf{u} , \mathbf{u}' , h to highest order

$$\begin{aligned} \mathbf{u}_{\text{pred}} &= \mathbf{u} + \Delta\tau \mathbf{u}' + \frac{\Delta\tau^2}{2} \mathbf{u}'' + \frac{\Delta\tau^3}{6} \mathbf{u}''' + \frac{\Delta\tau^4}{24} \mathbf{u}^{\langle 4 \rangle} + \frac{\Delta\tau^5}{120} \mathbf{u}^{\langle 5 \rangle}, \\ \mathbf{u}'_{\text{pred}} &= \mathbf{u}' + \Delta\tau \mathbf{u}'' + \frac{\Delta\tau^2}{2} \mathbf{u}''' + \frac{\Delta\tau^3}{6} \mathbf{u}^{\langle 4 \rangle} + \frac{\Delta\tau^4}{24} \mathbf{u}^{\langle 5 \rangle}, \\ h_{\text{pred}} &= h + \Delta\tau h' + \frac{\Delta\tau^2}{2} h'' + \frac{\Delta\tau^3}{6} h''' + \frac{\Delta\tau^4}{24} h^{(4)}. \end{aligned}$$

Force evaluation

$$\begin{aligned} \mathbf{P} &= \mathcal{L}_{\mathbf{u}}^T \mathbf{F}, \\ \mathbf{P}' &= \mathcal{L}_{\mathbf{u}'}^T \mathbf{F} + r \mathcal{L}_{\mathbf{u}}^T \dot{\mathbf{F}}. \end{aligned} \tag{2}$$

Derivatives for corrector

$$\begin{aligned} \mathbf{u}'' &= \frac{1}{2} (h \mathbf{u} + r \mathbf{P}), \\ \mathbf{u}''' &= \frac{1}{2} (h' \mathbf{u} + h \mathbf{u}' + r' \mathbf{P} + r \mathbf{P}'), \\ h' &= 2(\mathbf{u}' \cdot \mathbf{P}), \\ h'' &= 2(\mathbf{u}'' \cdot \mathbf{P} + \mathbf{u}' \cdot \mathbf{P}'). \end{aligned} \tag{3}$$

Fourth-order corrector (Hut, Makino & McMillan 1995)

$$\begin{aligned}
 \mathbf{u}_{\text{corr}} &= \mathbf{u}_{\text{old}} + \frac{\Delta\tau}{2}(\mathbf{u}'_{\text{corr}} + \mathbf{u}'_{\text{old}}) - \frac{\Delta\tau^2}{12}(\mathbf{u}''_{\text{new}} - \mathbf{u}''_{\text{old}}), \\
 \mathbf{u}'_{\text{corr}} &= \mathbf{u}'_{\text{old}} + \frac{\Delta\tau}{2}(\mathbf{u}''_{\text{new}} + \mathbf{u}''_{\text{old}}) - \frac{\Delta\tau^2}{12}(\mathbf{u}'''_{\text{new}} - \mathbf{u}'''_{\text{old}}), \\
 h_{\text{corr}} &= h_{\text{old}} + \frac{\Delta\tau}{2}(h'_{\text{new}} + h'_{\text{old}}) - \frac{\Delta\tau^2}{12}(h''_{\text{new}} - h''_{\text{old}}). \quad (4)
 \end{aligned}$$

Velocity derivatives for prediction

$$\begin{aligned}
 \mathbf{u}_{\text{mid}}^{\langle 4 \rangle} &= \frac{\mathbf{u}'''_{\text{new}} - \mathbf{u}'''_{\text{old}}}{\Delta\tau}, \\
 \mathbf{u}_{\text{mid}}^{\langle 5 \rangle} &= \frac{12}{\Delta\tau^2} \left(\frac{\mathbf{u}'''_{\text{new}} + \mathbf{u}'''_{\text{old}}}{2} - \frac{\mathbf{u}''_{\text{new}} - \mathbf{u}''_{\text{old}}}{\Delta\tau} \right). \quad (5)
 \end{aligned}$$

Shifted from midpoint

$$\begin{aligned}
 \mathbf{u}_{\text{new}}^{\langle 4 \rangle} &= \mathbf{u}_{\text{mid}}^{\langle 4 \rangle} + \frac{\Delta\tau}{2} \mathbf{u}_{\text{mid}}^{\langle 5 \rangle}, \\
 \mathbf{u}_{\text{new}}^{\langle 5 \rangle} &= \mathbf{u}_{\text{mid}}^{\langle 5 \rangle}, \\
 h'''_{\text{new}} &= h'''_{\text{mid}} + \frac{\Delta\tau}{2} h''''_{\text{mid}}, \\
 h''''_{\text{new}} &= h''''_{\text{mid}}.
 \end{aligned} \quad (6)$$

Taylor series for physical time

$$\begin{aligned}t' &= \mathbf{u} \cdot \mathbf{u}, \\t'' &= 2(\mathbf{u} \cdot \mathbf{u}'), \\t''' &= 2(\mathbf{u} \cdot \mathbf{u}'' + \mathbf{u}' \cdot \mathbf{u}'), \\t^{(4)} &= 2(\mathbf{u} \cdot \mathbf{u}''') + 6(\mathbf{u}' \cdot \mathbf{u}''), \\t^{(5)} &= 2(\mathbf{u} \cdot \mathbf{u}^{(4)}) + 8(\mathbf{u}' \cdot \mathbf{u}''') + 6(\mathbf{u}'' \cdot \mathbf{u}''), \\t^{(6)} &= 2(\mathbf{u} \cdot \mathbf{u}^{(5)}) + 10(\mathbf{u}' \cdot \mathbf{u}^{(4)}) + 20(\mathbf{u}'' \cdot \mathbf{u}''').\end{aligned}\quad (7)$$

Next look-up time

$$t_{\text{next}} = t + \sum_{n=1}^6 \frac{t^{(n)} \Delta\tau^n}{n!}.\quad (8)$$

Regularized time-step by inversion

$$-\Delta t + \sum_{n=1}^6 \frac{t^{(n)} \Delta\tau^n}{n!} = 0.\quad (9)$$

Newton-Raphson iteration to solve for $\Delta\tau$

$$x_{i+1} = x_i - \frac{f(x_i)}{f'(x_i)}.$$

Initial guess

$$\Delta\tau_1 = \frac{\Delta t_{bs}}{\Delta t} \Delta\tau_{orig},$$

with quantized time-step $\Delta t_{bs}(= 2^{-n} \leq \Delta t_{orig})$.

Commensurability condition

$$\text{mod}(t_{sys}, \Delta t_{bs}) = 0.$$

Solve equation (9) to obtain $\Delta\tau_{bs}$.

Three-Body Regularization

Initial conditions $\mathbf{r}_i, \mathbf{p}_i, \mathbf{p}_3 = -(\mathbf{p}_1 + \mathbf{p}_2)$

Basic Hamiltonian $\mu_{k3} = \frac{m_k m_3}{m_k + m_3}$

$$H = \sum_{k=1}^2 \frac{1}{2\mu_{k3}} \mathbf{p}_k^2 + \frac{1}{m_3} \mathbf{p}_1 \cdot \mathbf{p}_2 - \frac{m_1 m_3}{R_1} - \frac{m_2 m_3}{R_2} - \frac{m_1 m_2}{R}$$

KS coordinate transformation $\mathbf{Q}_k^2 = R_k, \quad (k = 1, 2)$

Time transformation $dt = R_1 R_2 d\tau$

Regularized Hamiltonian $\Gamma^* = R_1 R_2 (H - E_0)$

$$\begin{aligned} \Gamma^* = & \sum_{k=1}^2 \frac{1}{8\mu_{k3}} R_{3-k} \mathbf{P}_k^2 + \frac{1}{16m_3} \mathbf{P}_1^T \mathbf{A}_1 \cdot \mathbf{A}_2^T \mathbf{P}_2 \\ & - m_1 m_3 R_2 - m_2 m_3 R_1 - \frac{m_1 m_2 R_1 R_2}{|\mathbf{R}_1 - \mathbf{R}_2|} - E_0 R_1 R_2 \end{aligned}$$

Equations of motion

$$\frac{d\mathbf{Q}_k}{d\tau} = \frac{\partial \Gamma^*}{\partial \mathbf{P}_k}, \quad \frac{d\mathbf{P}_k}{d\tau} = -\frac{\partial \Gamma^*}{\partial \mathbf{Q}_k}$$

Regular solutions: $R_1 \rightarrow 0$ or $R_2 \rightarrow 0$

Singular term < regular terms: $|\mathbf{R}_1 - \mathbf{R}_2| > \max(R_1, R_2)$

Three-Body Transformations

Coordinates & momenta $\mathbf{q}_k = \tilde{\mathbf{q}}_k - \tilde{\mathbf{q}}_3, \quad \mathbf{p}_k = \tilde{\mathbf{p}}_k$

Regularized coordinates ($q_1 \geq 0$)

$$Q_1 = [\tfrac{1}{2}(|\mathbf{q}_1| + q_1)]^{1/2}$$

$$Q_2 = \tfrac{1}{2}q_2/Q_1$$

$$Q_3 = \tfrac{1}{2}q_3/Q_1$$

$$Q_4 = 0$$

Regularized momenta $\mathbf{P}_k = \mathbf{A}_k \mathbf{p}_k$

Basic matrix

$$\mathbf{A}_1 = 2 \begin{bmatrix} Q_1 & Q_2 & Q_3 & Q_4 \\ -Q_2 & Q_1 & Q_4 & -Q_3 \\ -Q_3 & -Q_4 & Q_1 & Q_2 \\ Q_4 & -Q_3 & Q_2 & -Q_1 \end{bmatrix}$$

KS transformations $\mathbf{q}_k = \tfrac{1}{2}\mathbf{A}_k^T \mathbf{Q}_k$

Physical momenta $\mathbf{p}_k = \tfrac{1}{4}\mathbf{A}_k^T \mathbf{P}_k / R_k$

Coordinates & momenta

$$\tilde{\mathbf{q}}_3 = -\sum_{k=1}^2 m_k \mathbf{q}_k / M$$

$$\tilde{\mathbf{q}}_k = \tilde{\mathbf{q}}_3 + \mathbf{q}_k$$

$$\tilde{\mathbf{p}}_k = \mathbf{p}_k$$

$$\tilde{\mathbf{p}}_3 = -(\mathbf{p}_1 + \mathbf{p}_2) \quad (k = 1, 2)$$

PN Codes

Two-body Hermite KS toy code

alternative PN formulations

Three-body AZ regularization

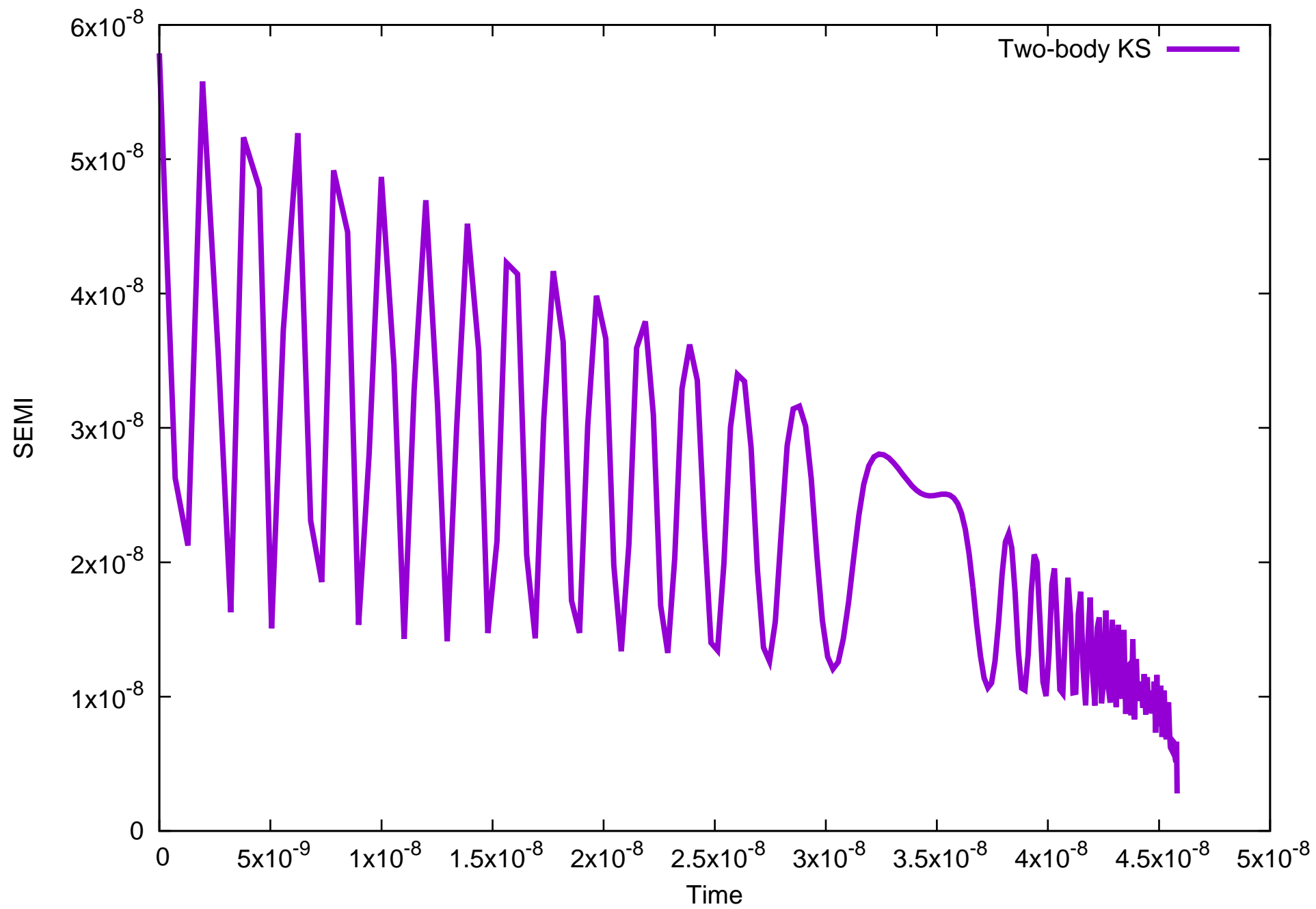
dominant binary, sequential terms

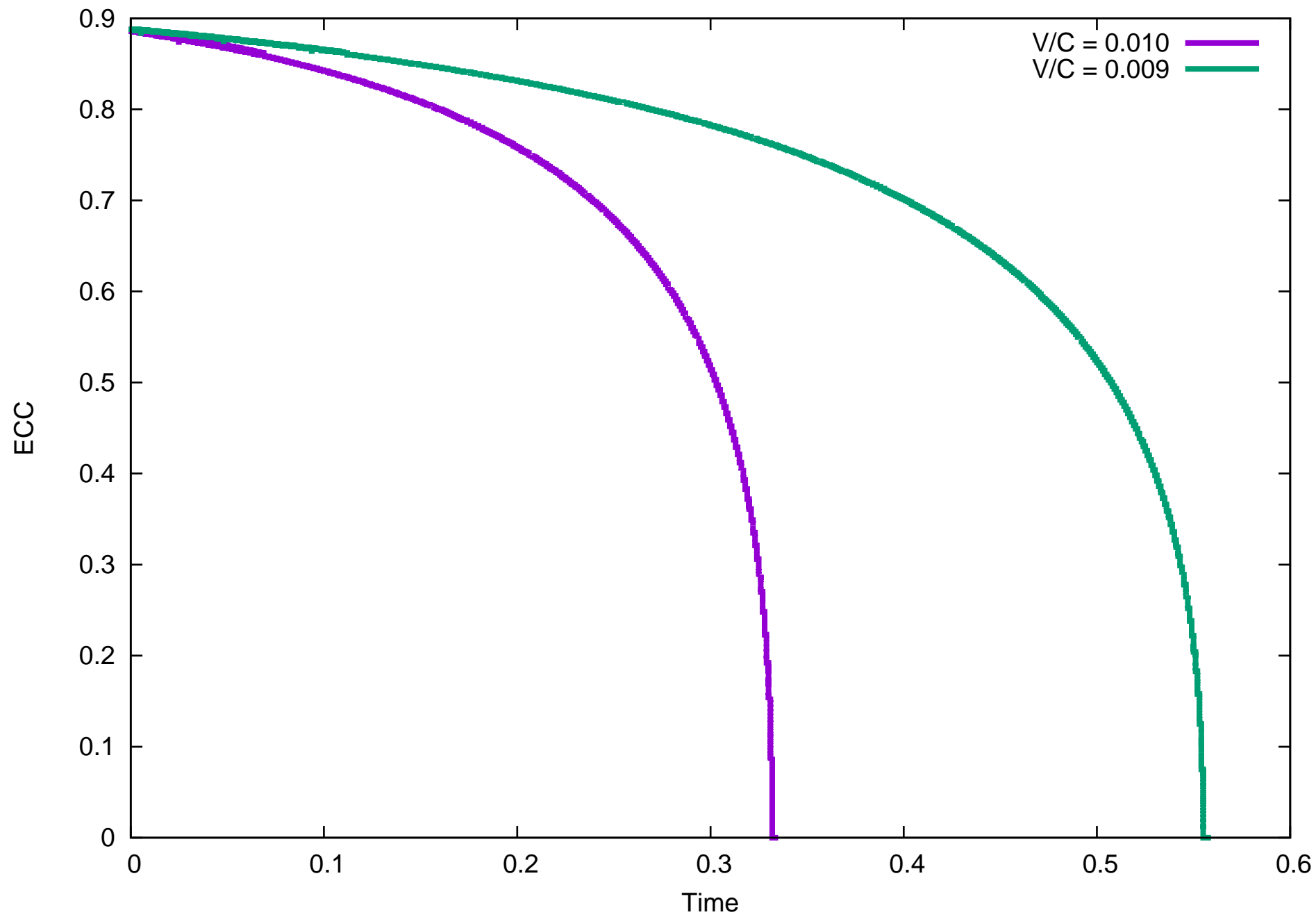
homework: Einstein shift

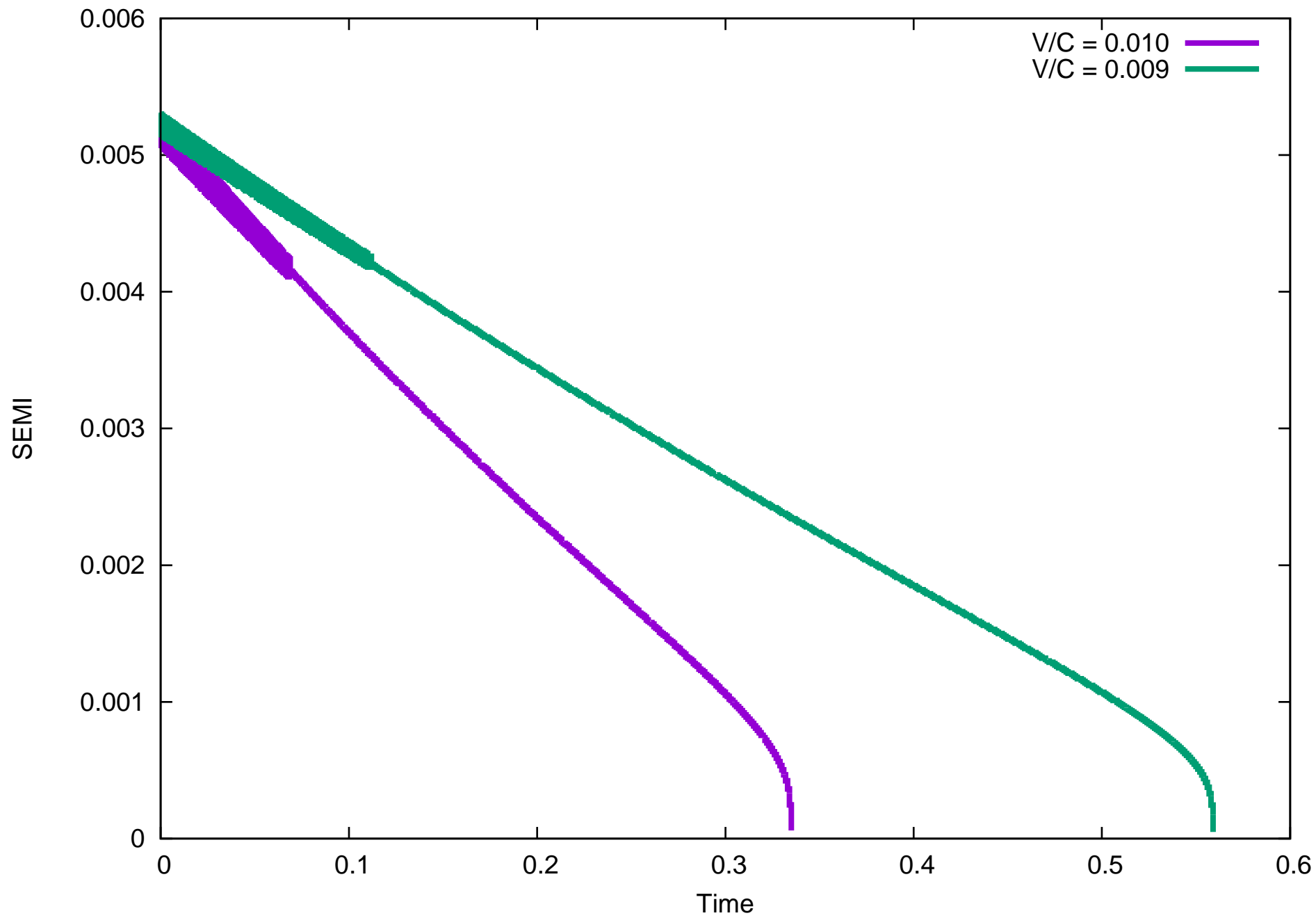
Chain Algorithmic regularization

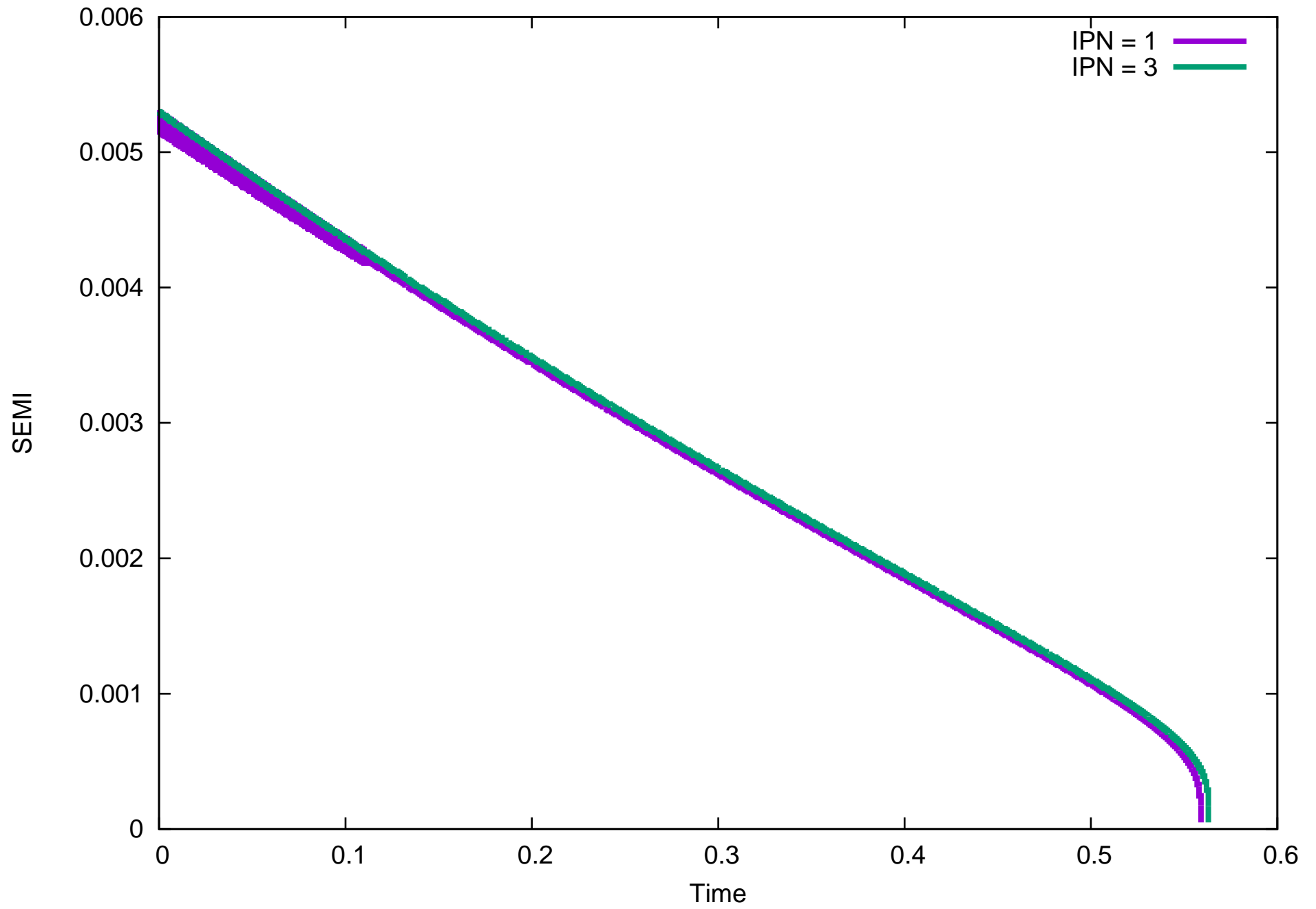
compact sub-system

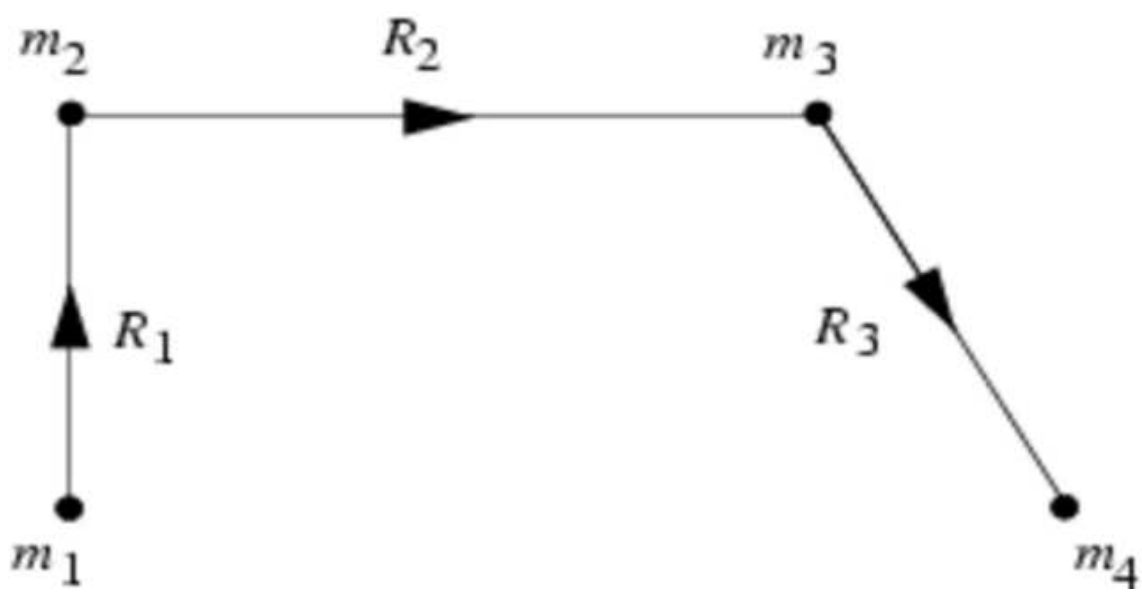
sequential PN, unperturbed KS











Initialization

Two-body $R < R_{cl}, \dot{R} < 0, \Rightarrow$ KS

Hierarchy $B + S \Rightarrow H,$ Stability test

Triple chain $a_1(1 - e_1) < 3a_0, \dot{r} < 0, \Rightarrow$ T

Four-body chain $B + B, \dot{r} < 0, \Rightarrow$ B-B

Higher orders $T + B, \dot{r} < 0, \Rightarrow$ Q

Chain Initialization

Unperturbed exit	$\Delta\omega > 1 \times 10^{-3}$
Activate chain	$N_{\text{ch}} = 2$ or > 2
Alternative path	perturbed KS \Rightarrow search
Selection	chain members & c.m.
Initialization	c.m. force & chain vectors
Synchronization	block time-step
Interface	perturber list

Termination

Hyperbolic KS $R > R_0$ or $\gamma > 0.001$

Binary KS $\gamma > 0.1$, $\gamma = \frac{(\mathbf{F}_1 - \mathbf{F}_2)R^2}{m_1 + m_2}$

Collision $KS \Rightarrow S$

Triple $d_3 > 20a$, $\Rightarrow B + S$

Hierarchy $a_1(1 - e_1) < Sa_0$, $\gamma > 0.01$

ARchain $R_1 \ll R_2 + R_3 \Rightarrow B + 2S$

Relativistic Effects

Einstein shift

$$\Delta\omega = \frac{6\pi M}{ac^2(1 - e^2)}$$

GR radiation time-scale

$$t_{\text{GR}} = \frac{5}{64} \frac{c^5 g(e) a^4}{X(1 + X) m_1^3}$$

$$X = \frac{m_2}{m_1}, \quad g(e) \simeq \frac{(1 - e^2)^{7/2}}{4.5}$$

Speed of light & Schwarzschild radius

$$c = \frac{3 \times 10^5}{V^*}, \quad R_{\text{Sch}} = \frac{2M}{c^2}$$

Velocities

$$V^* = \alpha \left(\frac{M}{R_{pc}} \right)^{1/2}, \quad v_{\text{max}} = \left(\frac{M}{a} \right)^{1/2} \left(\frac{1 + e}{1 - e} \right)^{1/2}$$

PN conditions

$$\Delta\omega > 1 \times 10^{-4}, \quad t_{\text{GR}} < t^*, \quad r < n R_{\text{Sch}}$$

Post-Newtonian Terms

Equation of motion $\frac{d^2 \mathbf{r}}{dt^2} = \frac{M}{r^2} \left[(-1 + A) \frac{\mathbf{r}}{r} + B \mathbf{v} \right]$

First-order precession $M = m_1 + m_2, \quad \eta = \frac{m_1 m_2}{M^2}$

$$A_1 = 2(2 + \eta) \frac{M}{r} - (1 + 3\eta)v^2 + \frac{3}{2}\eta \dot{r}^2$$

$$B_1 = 2(2 - \eta)\dot{r}$$

Higher-order precession $A_2 = \dots, \quad B_2 = \dots, \quad A_3 = \dots, \quad B_3 = \dots$

Gravitational radiation $A_{5/2} = \frac{8}{5}\eta \frac{M}{r} \dot{r} \left(\frac{17M}{3r} + 3v^2 \right)$

$$B_{5/2} = -\frac{8}{5}\eta \frac{M}{r} \left(3\frac{M}{r} + v^2 \right)$$

Total GR perturbation

$$\mathbf{P}_{GR} = \frac{M}{c^2 r^2} \left[\left(A_1 + \frac{A_2}{c^2} + \frac{A_{5/2}}{c^3} \right) \frac{\mathbf{r}}{r} + \left(B_1 + \frac{B_2}{c^2} + \frac{B_{5/2}}{c^3} \right) \mathbf{v} \right]$$

GR spin dominant BH

ARC energy change $\Delta E = \frac{m_1 m_2}{M} \int (\mathbf{P}_{GR} + \mathbf{P}_{dyn}) \cdot \mathbf{v} dt$

Decision-Making

Increasing GR effect	$t_{\text{GR}} < 500, 50, 1$
	IPN = 1, 2, 3
GR time-scale	$T_z = \frac{150a^4}{M^3}(1 - e^2)^{7/2} \text{ Myr}$
Graduated PN	$\Delta w > (1, 10, 100) \times 10^{-4} \Rightarrow \text{IPN} = 1, 2, 3$
Coalescence	$R < 4R_{\text{Sch}} = \frac{8M}{c^2}$
Alternative merging	IPN = 3, $N = 2$, $N_p = 0$
	IPN ≥ 2 , $a(1 - e) < R_{\text{Sch}}$
	IPN = 3, $N = 3$, $a_1(1 - e_1) > 100a$
Unperturbed KS	$t_{\text{GR}} < 500$ & $\Delta w > 1 \times 10^{-4}$, $\gamma < 1 \times 10^{-6}$
Tidal disruption	KS or CHAIN

PN Elements

Energy $\epsilon_b = \epsilon_0 + \frac{\epsilon_1}{c^2} + \frac{\epsilon_2}{c^4} + \frac{\epsilon_3}{c^6}, \quad a = -\frac{M}{2\epsilon_b}$

$$\epsilon_0 = \frac{1}{2}V^2 - \frac{M}{R}, \quad \eta = \frac{m_1 m_2}{M^2}$$

$$\epsilon_1 = \frac{1}{2} \frac{M}{R} + \frac{3}{8}(1-3\eta)V^4 + \frac{1}{2} \left((3+\eta)V^2 + \eta\dot{R}^2 \right) \frac{M}{R}$$

Angular momentum $\mathbf{J} = \mathbf{J}_0(1 + f_1/c^2 + f_2/c^4)$

Eccentricity $e^2 = \left(1 - \frac{\mathbf{J}^2}{Ma}\right)$

Lenz vector $\mathbf{e} = \mathbf{V} \times \mathbf{R} \times \mathbf{V}/M - \mathbf{R}/R$

Periapse advance $\Delta\omega = \frac{6\pi M}{c^2 a(1-e^2)}$

Kozai-Lidov cycles

$$T_{\text{Kozai}} = \frac{T_{\text{out}}^2}{T_{\text{in}}} \left(\frac{1+q_{\text{out}}}{q_{\text{out}}} \right) (1-e_{\text{out}}^2)^{3/2} g(e_{\text{in}}, \omega_{\text{in}}, \psi)$$

Unperturbed GR Orbit

Compact objects $\max (K_1^*, K_2^*) > 12$

Derivatives $\dot{a} = -\frac{64M_1M_2(M_1 + M_2)}{5c^5a^3(1 - e^2)^{7/2}}$

$$\dot{e} = -\frac{304eM_1M_2(M_1 + M_2)}{15c^5a^4(1 - e^2)^{5/2}}g(e)$$

Einstein shift $\Delta\omega = \frac{6\pi(M_1 + M_2)}{ac^2(1 - e^2)}$

KS rotation $\theta = \frac{\Delta\omega\Delta t}{T_K}$, rotate by $\theta/2$

New elements $a_1 = a + \dot{a}\Delta t$, $e_1 = \max (e + \dot{e}\Delta t, 0)$

Energy updating $\Delta E = \mu(H_{\text{old}} - H_{\text{new}})$

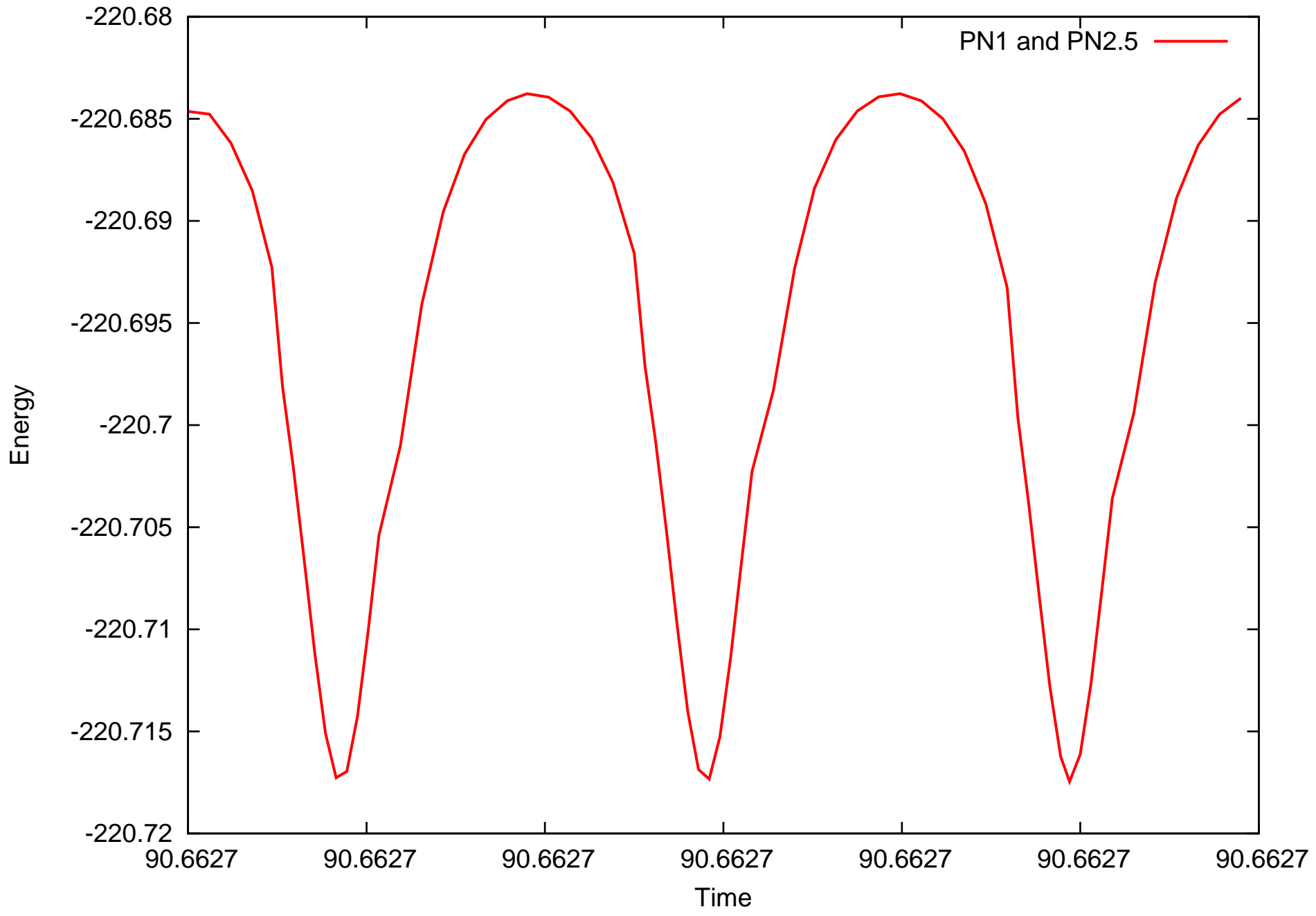
Modify KS orbit Shrink, $e = \text{const}$; decrease, $H = \text{const}$

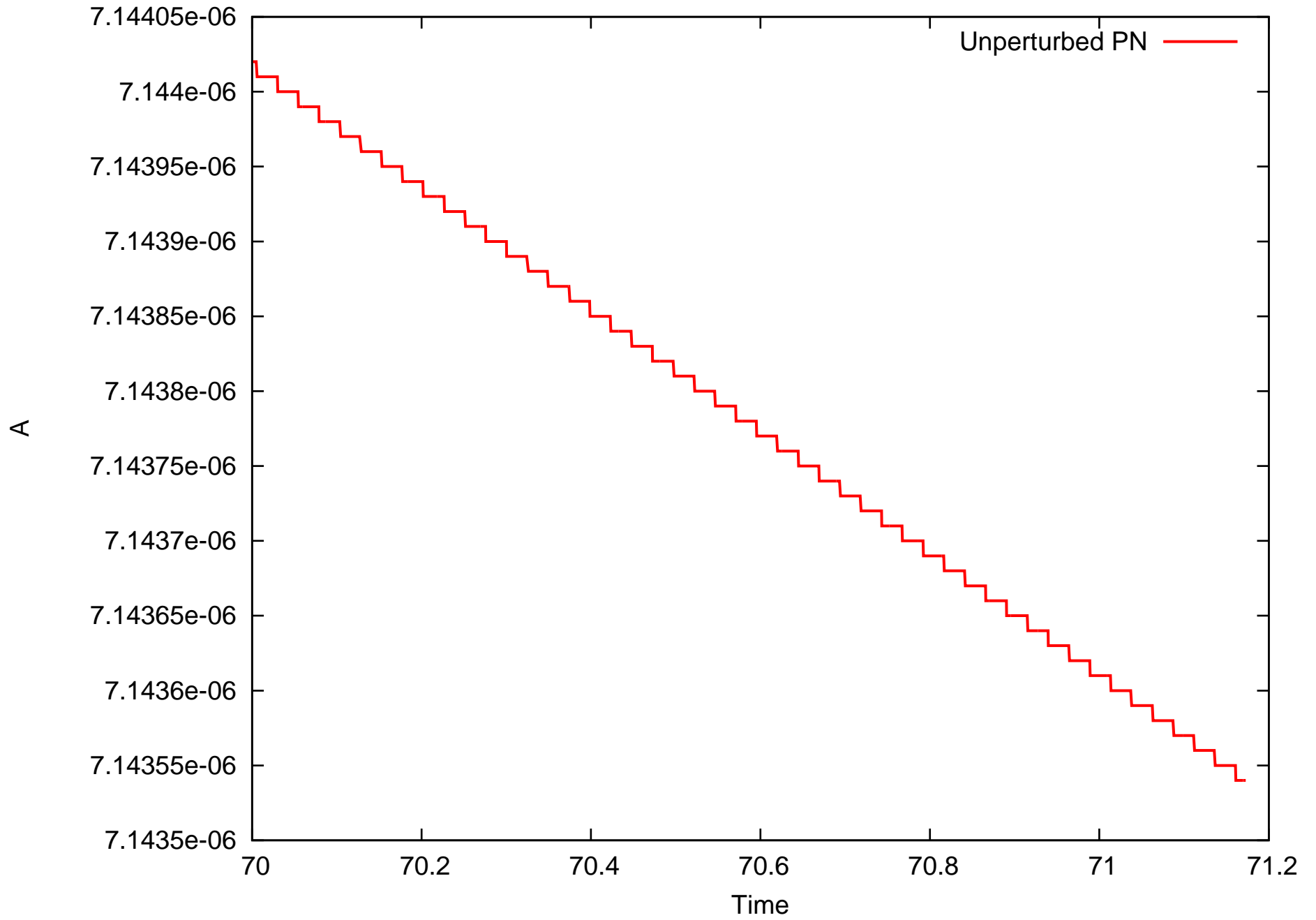
Validity $1 \times 10^{-4} < \Delta\omega < 1 \times 10^{-3}$

Unperturbed KS $\gamma < 1 \times 10^{-6}$

Membership

Reduction	escape R_1 or $R_{n-1} > R^*$, $\dot{r} > 0$
Double KS	$n = 4$, two well separated pairs
Injection	$\gamma > 1 \times 10^{-2}$ & $r_p < R^*$, $\dot{r}_p < 0$
Retention	$n = 2$, $IPN > 0$, $\Delta\omega > 1 \times 10^{-3}$
Perturbers	massive binary, $N_p \simeq 3 - 10$
Perturbation	$\gamma \simeq \frac{2m_p}{M_{bh}} \max\left(\frac{R_k}{d_p}\right)^3$





Energy Corrections

Unperturbed KS $1 \times 10^{-4} < \Delta\omega < 1 \times 10^{-3}$

Perturbed KS $\mathbf{P}_{2.5}$ & $\dot{\mathbf{P}}_{2.5}$ added to \mathbf{F} & $\dot{\mathbf{F}}$

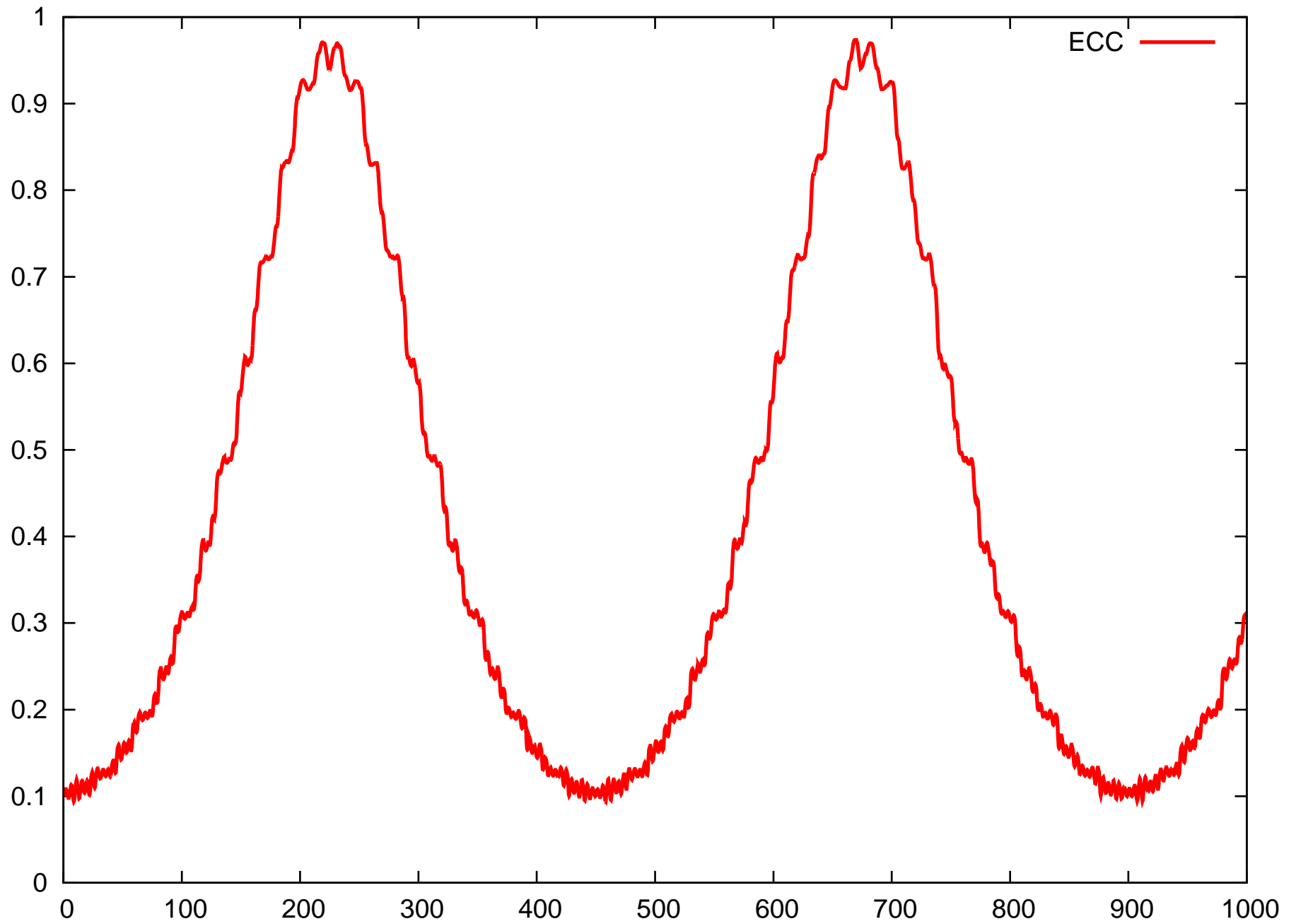
Energy derivatives $\dot{E} = \mathbf{P} \cdot \mathbf{v}$, $\ddot{E} = \mathbf{P} \cdot \dot{\mathbf{v}} + \dot{\mathbf{P}} \cdot \mathbf{v}$

Hermite method
$$\Delta h = \frac{1}{2}(\dot{E}_0 + \dot{E})\Delta t$$
$$+ \frac{1}{12}(\ddot{E}_0 - \ddot{E})\Delta t^2$$

Energy loss $\sum \mu \Delta h$

Conservation $E_{\text{ch}} = E_{\text{sub}} - \sum \Delta E$

$$E_{\text{tot}} = E_N + E_{\text{ch}}$$



Tidal Disruption

Classical condition $r_t = \left(\frac{m_{\text{BH}}}{m^*}\right)^{1/3} r^*$

Case A Two-body regularization

Pericentre test: $a(1 - e) < r_t$

Transform to apocentre $\theta = \pi/2$

Mass transfer $\Delta m = 0.1 m^*$

Escape velocity kick

Corrections KS elements and energy

Case B Chain regularization

Mass transfer $\Delta m = 0.1 m^*$

Ejection Energy corrections

Updating Chain membership and tables

Initialization Chain c.m. polynomials

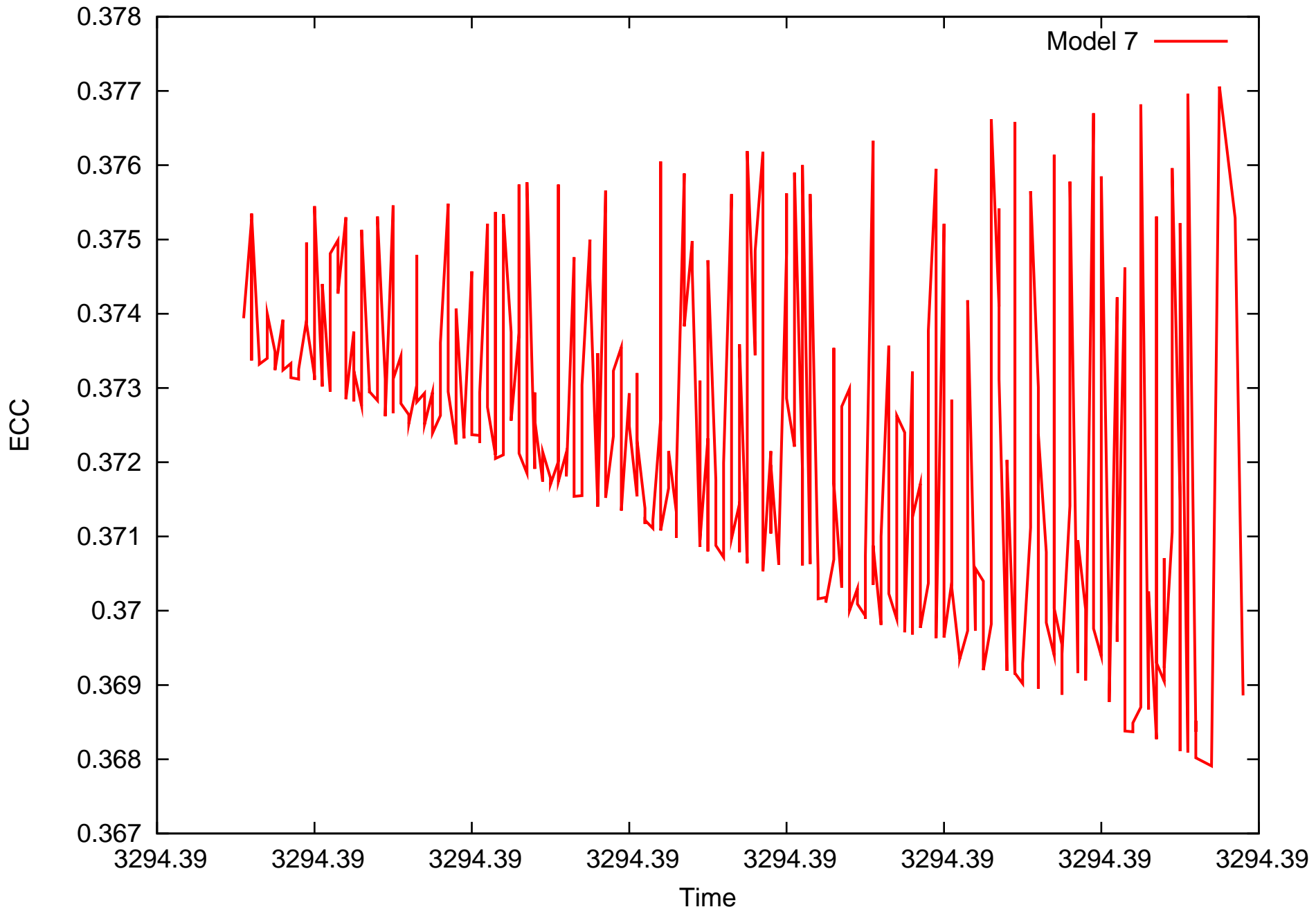
Requirements Large eccentricity

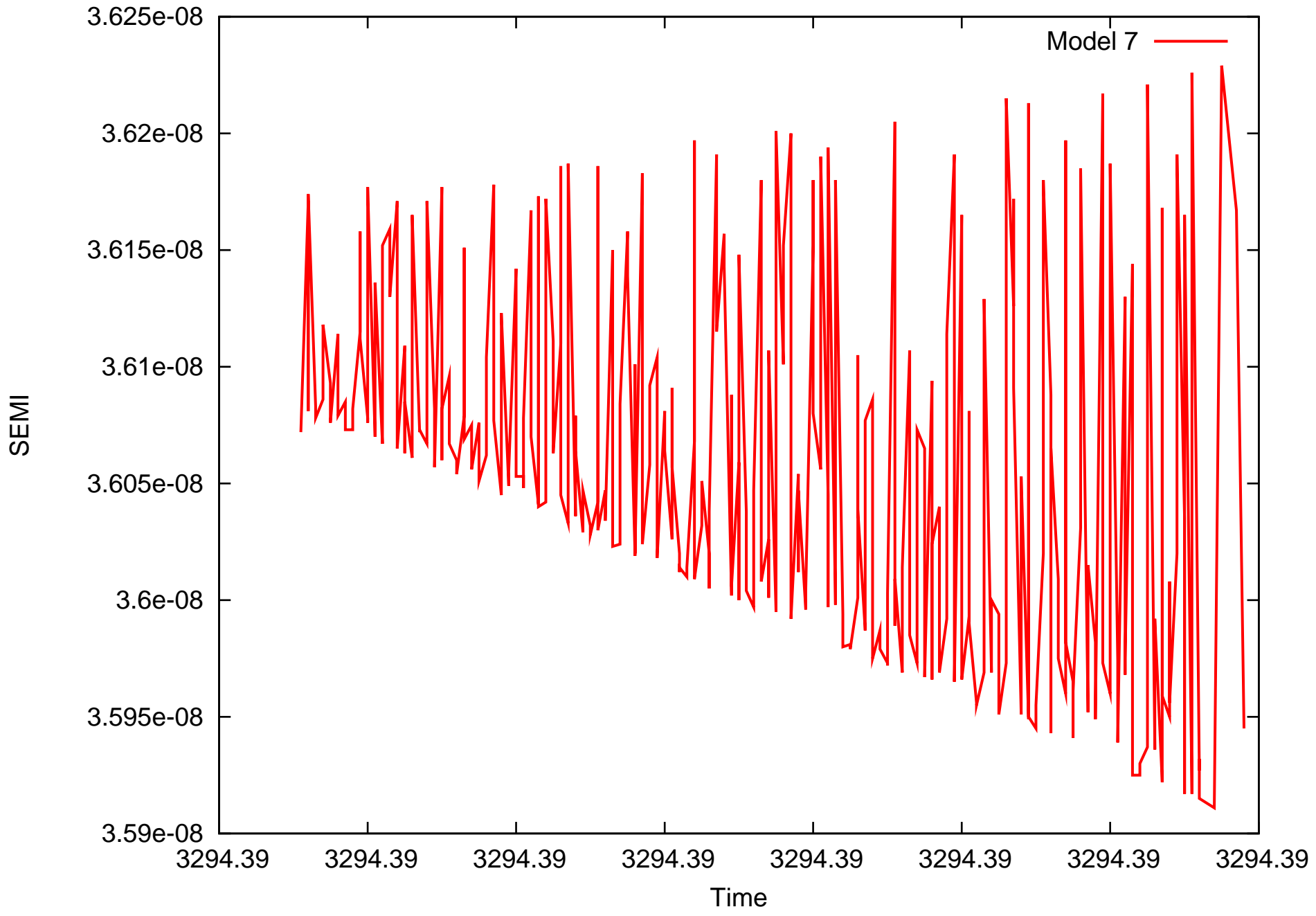
Effects Mass segregation & Kozai-Lidov

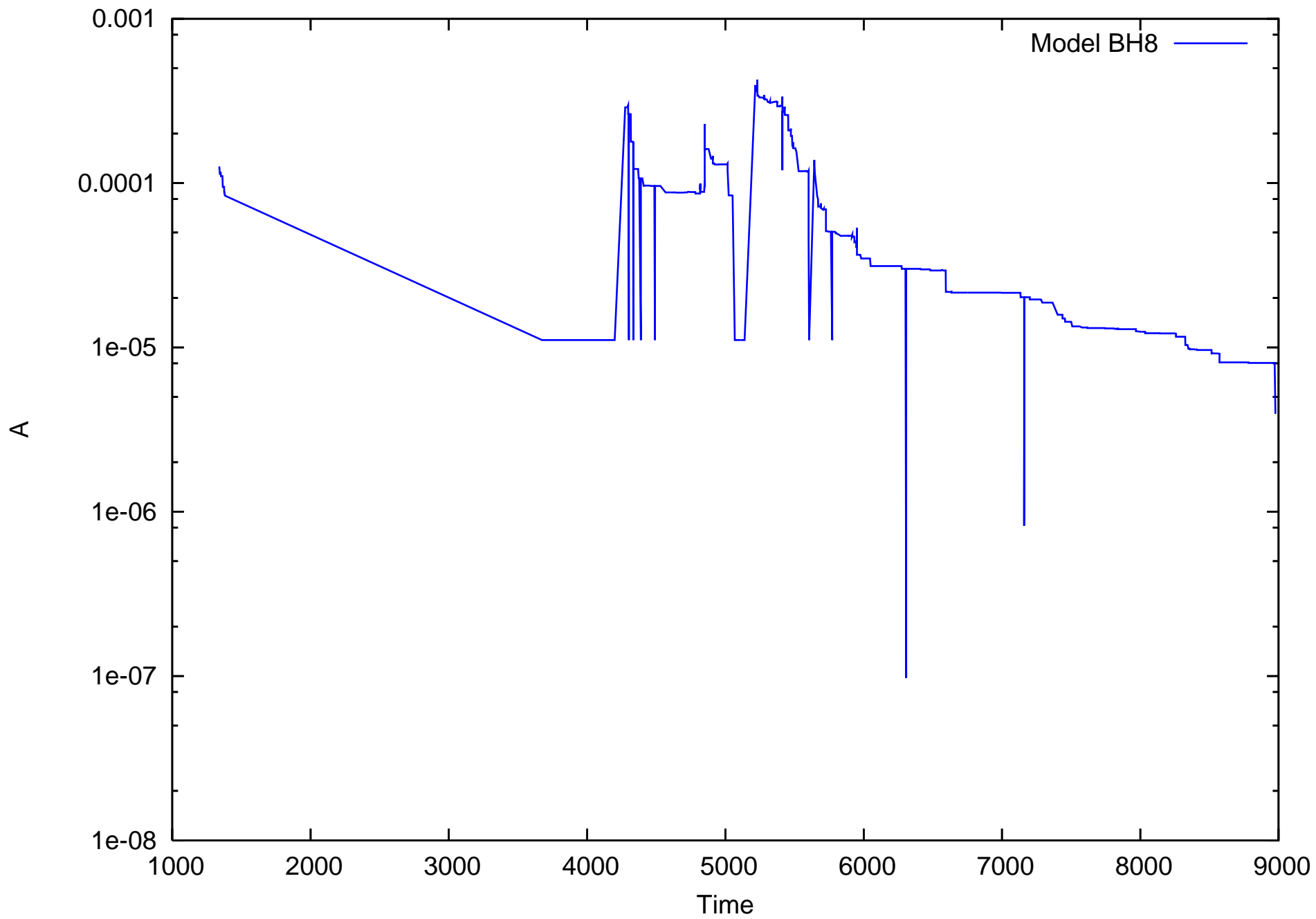
Special cases WD & AGB cores

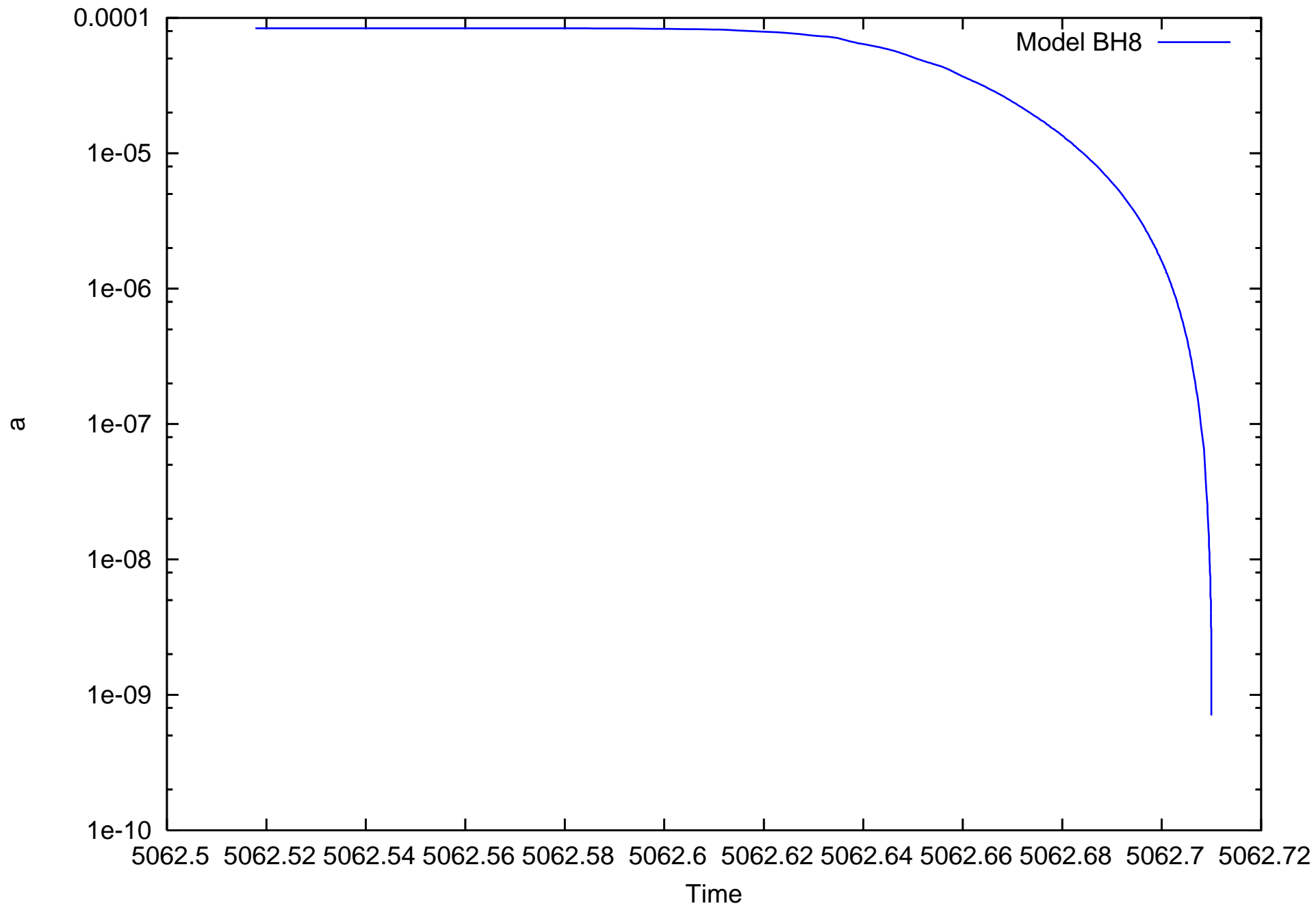
Astrophysics Instantaneous accretion

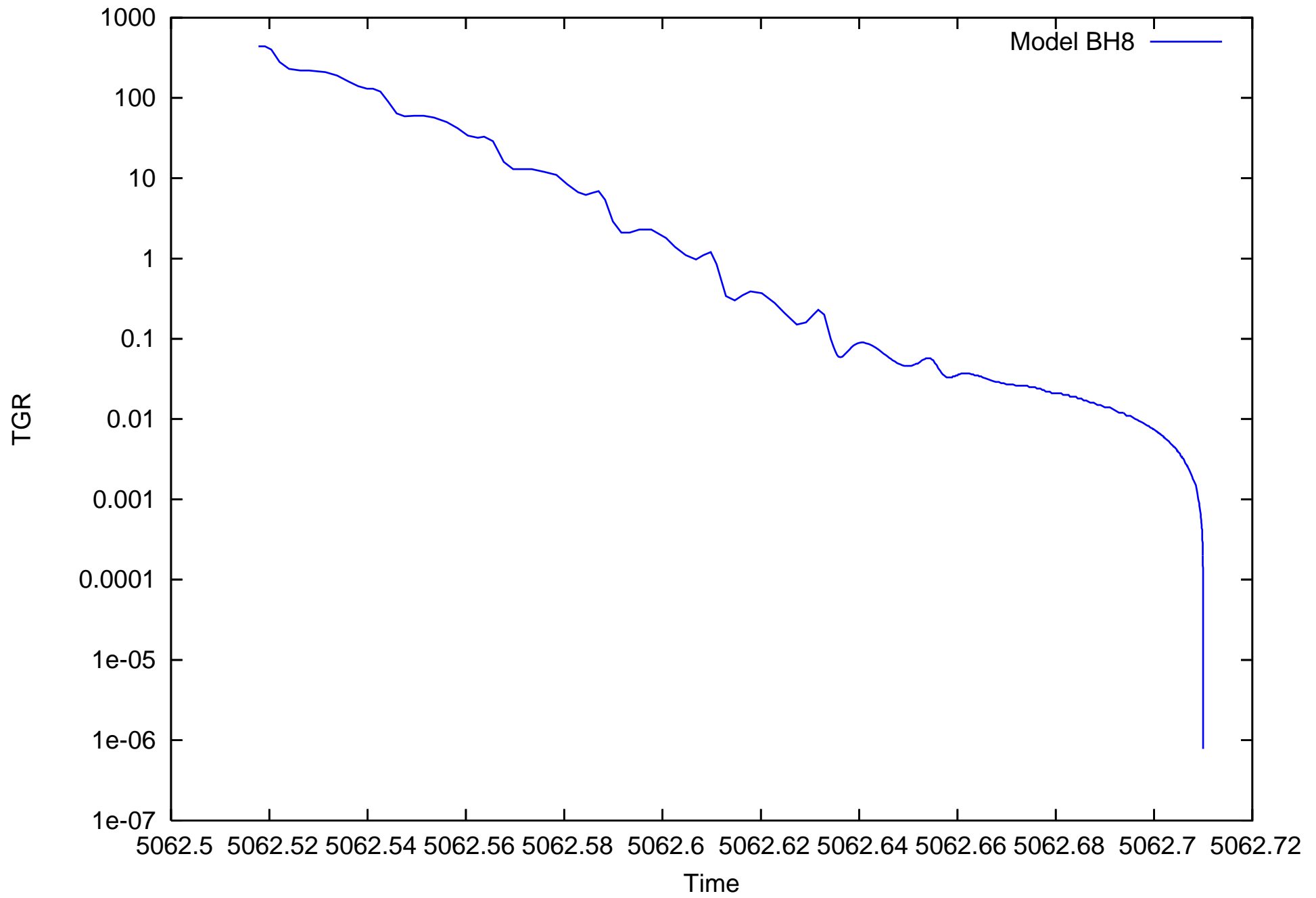
Observations Light flickering & emission lines











Discussion

IMF	upper limit, Z
Mass outcome	BH from BSE
PN simulations	speed of light $c = \frac{3 \times 10^5}{V^*}$
Dynamics	binary shrinkage & Kozai-Lidov
Slingshot	fast escapers
Retention	velocity kicks
Depletion	ejections of BH & BBH
Observations	BH + S binaries, hyper-velocities
Astropysics	tidal disruption & BH accretion