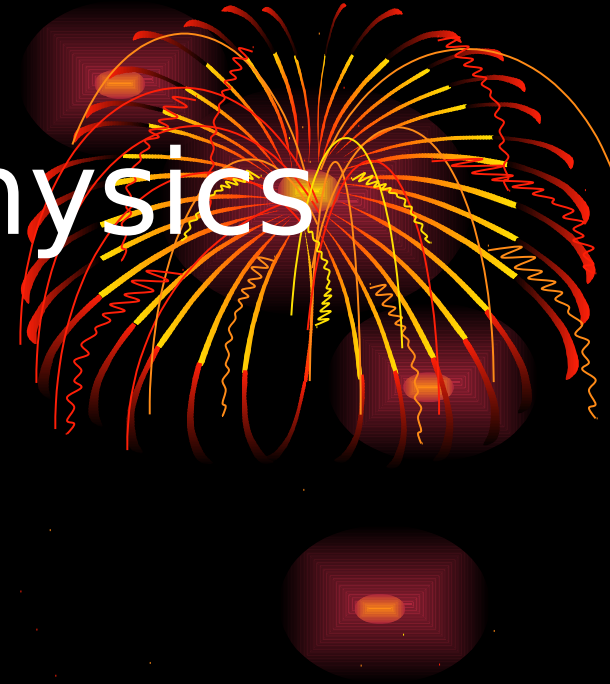


Computational Physics

Computerphysik



Rainer Spurzem, Astronomisches Rechen-Institut Zentrum
für Astronomie, Universität Heidelberg
Rüdiger Pakmor, Heidelberg Institute for Theoretical Studies

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Computer Physics

● Traditional:

Theory

Experiment/Observation

● Modern:

Theory



Experiment
Observation

Computer-Experiment

Checking Theories!

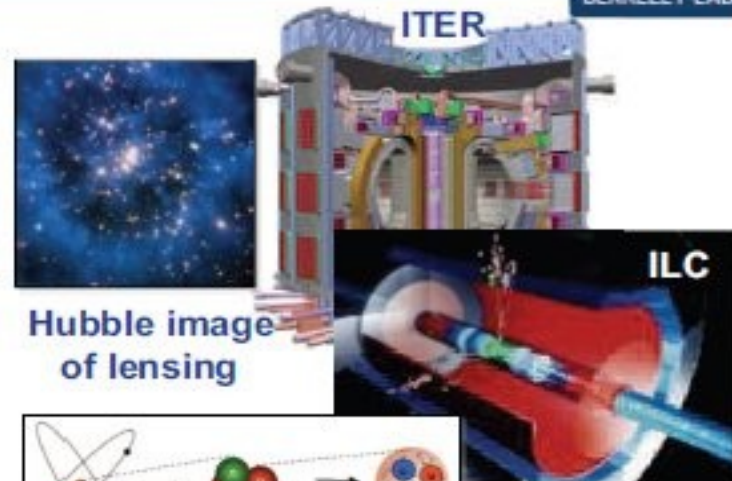
Predict Experiments/Observations!

Exascale simulation will enable fundamental advances in basic science

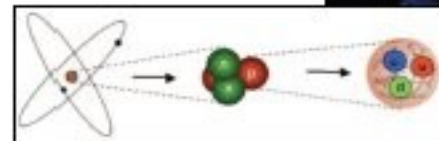


- High Energy & Nuclear Physics
 - Dark-energy and dark matter
 - Fundamentals of fission fusion reactions
- Facility and experimental design
 - Effective design of accelerators
 - Probes of dark energy and dark matter
 - ITER shot planning and device control
- Materials / Chemistry
 - Predictive multi-scale materials modeling: observation to control
 - Effective, commercial technologies in renewable energy, catalysts, batteries and combustion
- Life Sciences
 - Better biofuels
 - Sequence to structure to function

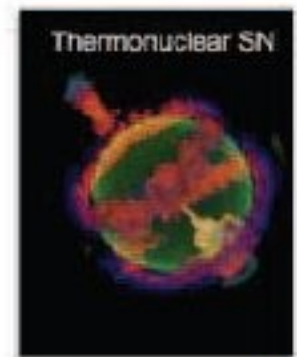
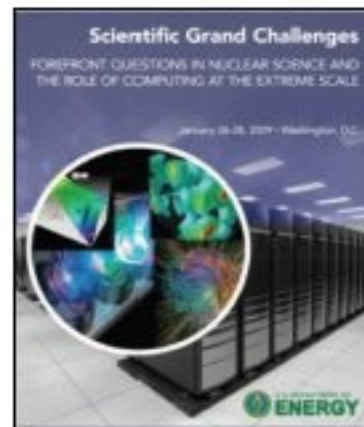
These breakthrough scientific discoveries and facilities require exascale applications and resources



Hubble image of lensing



Structure of nucleons



Advanced Computation in Energy Science at LBNL



Probe natural systems under constraints that are difficult or impossible to impose in the field or laboratory

Reveal the manner in which large-scale phenomena arise from smaller-scale properties

Discover new materials for green technology applications through first-principles calculations

Global Scale Reactive Transport Modeling of CH₄ hydrates (M. Reagan)



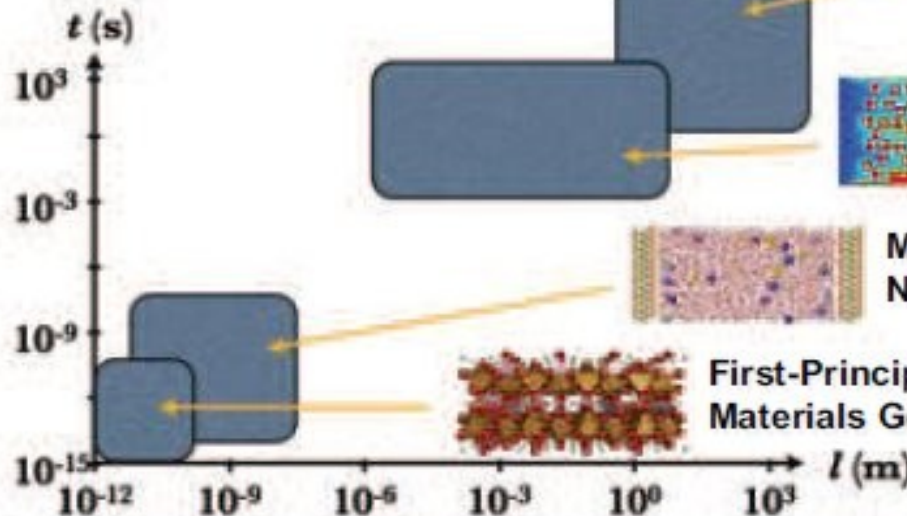
Pore Scale Reactive Transport Modeling of CO₂ sequestration (D. Trebotich)



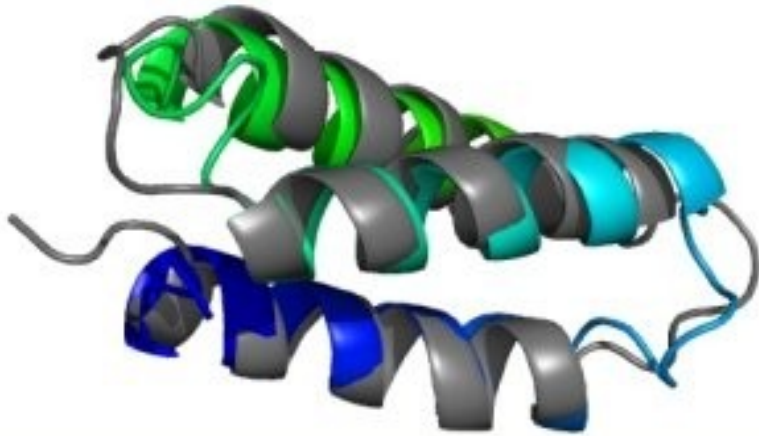
Molecular Dynamics Simulations of Natural Nanofluids (I. Bourg)



First-Principles Calculations of Materials Genome (K. Persson)



Research http://www.fz-juelich.de/ias/jsc/EN/Research/research_node.html
JSC's research and development concentrates on mathematical modelling and numerical, especially parallel algorithms for quantum chemistry, molecular dynamics and Monte-Carlo simulations. The focus in the computer sciences is on cluster computing, performance analysis of parallel programs, visualization, computational steering and grid computing.



Modelling and Simulation

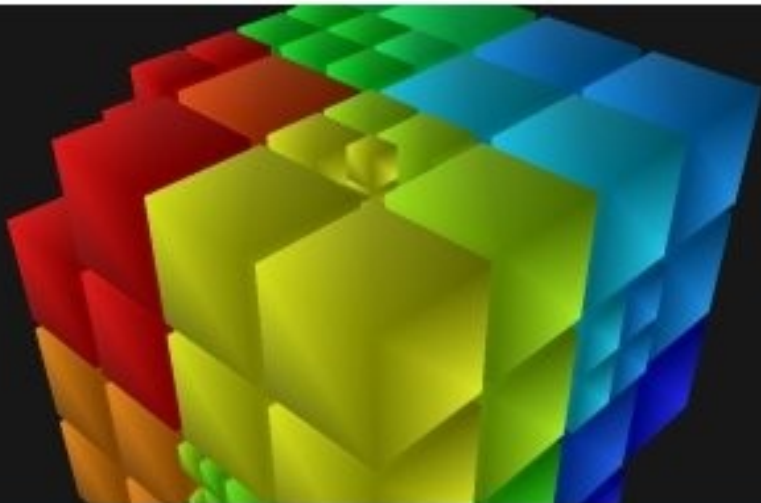
The simulation of complex systems in natural science or engineering depends on the development of adequate mathematical models. Thus the development of realistic and yet efficient models is a core activity at JSC. Examples of simulations are:

- Computational Plasma Physics
- ☞ Protein Folding
- Quantum Information Processing
- Civil Security and Traffic

Algorithms and Methods

Efficient simulations need powerful algorithms and methods. JSC focusses on the development of the following methods:

- Fast Coulomb Solvers
- Parallel-In-Time Integration
- Fast Multipole Method
- ☞ Parallel I/O



Why Computational Physics? Our view at this lecture...

- Differential Equations (ODE/PDE) – no analytical solution!

Relation to nonlinear dynamics, theory of chaos

- Ordinary differential equations (ODE) – initial value problems – particle simulations

Plasma physics, molekular dynamics, stellar dynamics...

- Ordinary differential equations (ODE) – boundary values problems – *quantum mechanics (time-indep.), statistical mechanics, stationary states*

- Partial differential equations (PDE) – initial and boundary value problems – mesh based or finite elements

fluid dynamics, numerical relativity, quantum mechanics (time-dep.)

- Monte Carlo methods – use of (pseudo) random numbers

numerical integration, stellar dynamics, particle physics (many dims.)

- Computational mathematics – “exact” mathematical computations

Mathematica, Maple, MatLab, ...

History

- Erik Holmberg (1908-2000)

Dissertation Univ. Lund (Sweden) (1937):

“A study of double and multiple galaxies”

Galaxies distributed in groups and pairs –

Satellite galaxies have uneven distribution

(Holmberg-Effect)



- **Father of numerical Astrophysics....**

- **...with 200 light bulbs**

History



The Astrophysical Journal, Nov. 1941

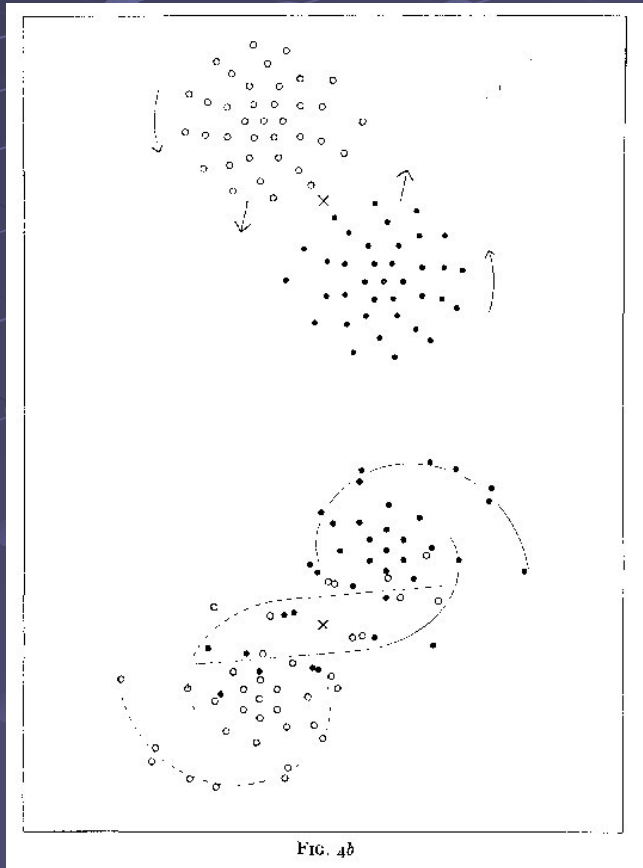


FIG. 4b

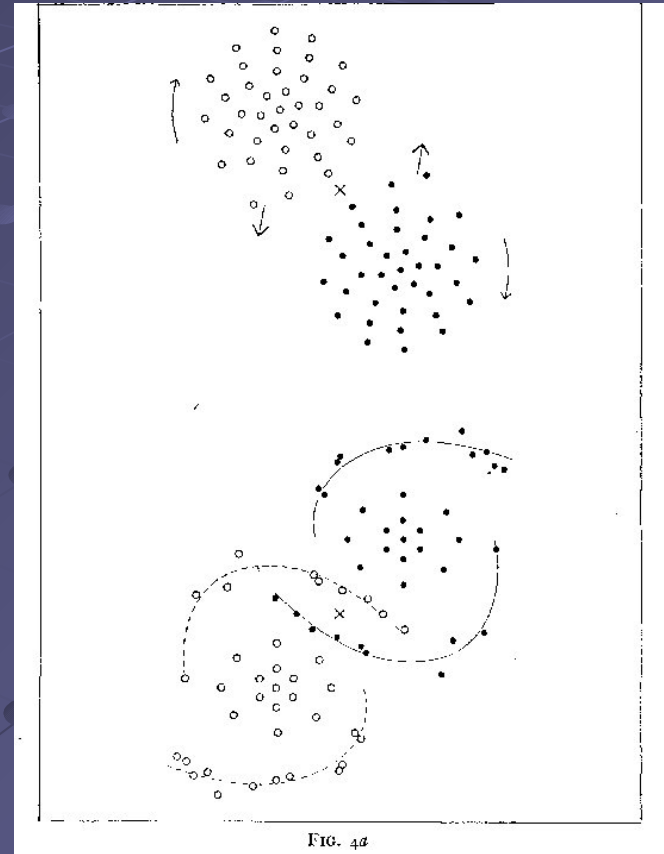


FIG. 4a

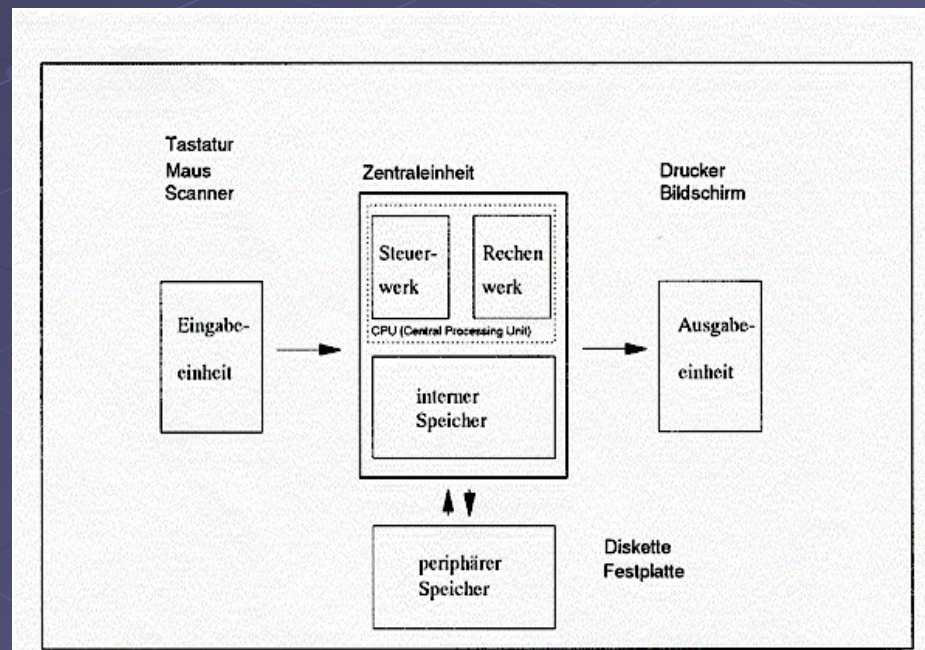
Geschichte

John von Neumann (1903-1957)

Geb. Budapest, Dozent Berlin,

ab 1930 Princeton Univ. NJ USA Princeton

„Requirements for an electronic computing machine“ (1946)



Introduction – Supercomputers

Supercomputer
JUGENE
IBM Blue Gene
Jülich
Supercomputing
Centre (JSC)
223 Tflop/s
...Petaflop/s...

2019: Exaflop/s System
(...Trillion...)

Opening with J. Rüttgers June 2008



Computational Science...

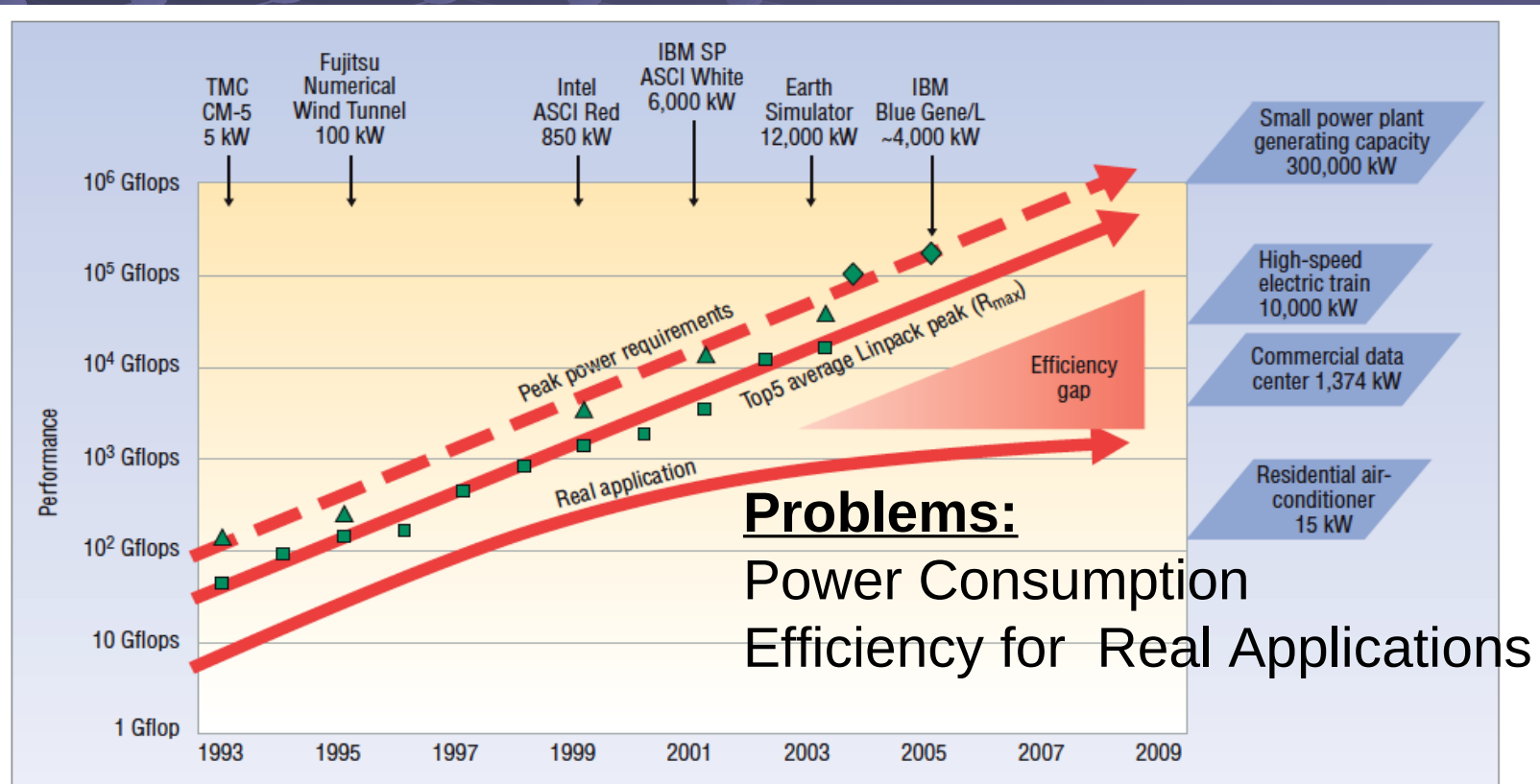
Exaflop/s?

...after von Neumann...

Petaflop/s

Teraflop/s

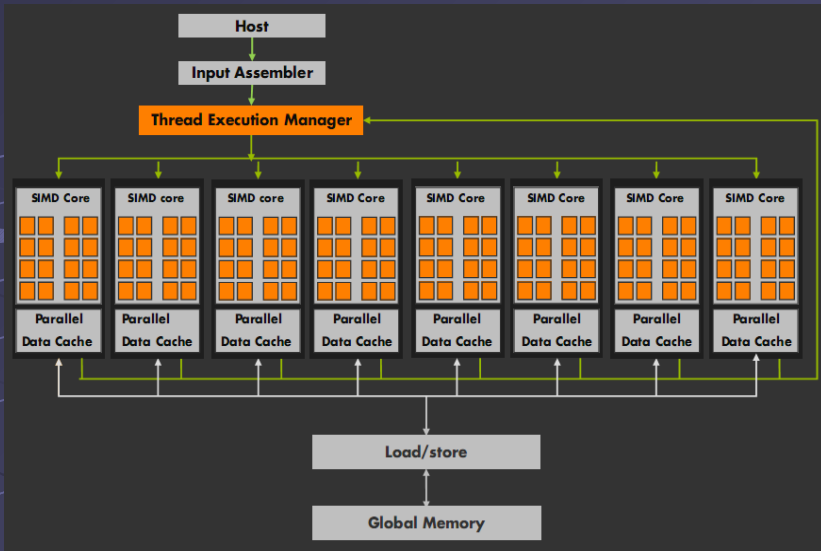
Gigaflop/s



Thanks to Horst Simon, LBNL/NERSC for this diagram.

Figure 1. Rising power requirements. Peak power consumption of the top supercomputers has steadily increased over the past 15 years.

Hardware



Each core

- 8 functional units
- SIMD 16/32 "warp"
- 8-10 stage pipeline
- Thread scheduler
- 128-512 threads/core
- 16 KB shared memory

Parallel Data Cache

Total #threads/chip
 $16 * 512 = 8K$

NVIDIA Tesla C1060 :

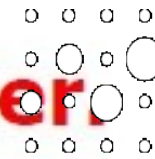
1.3 GHz * 240 processors * 3 flopclock = 930 Gflops

NAOC laohu cluster Beijing, China



Heidelberg

Kepler GPU cluster



VolkswagenStiftung

Kepler GPU cluster

12 nodes = 12 x 16 = 192 CPU cores (@ 2 GHz)

12 x 64 GB = 768 GB RAM CPU memory

12 GPUs K20m = 12 x 2496 ~ 30k GPU threads

12 x 4.8 GB ~ 57 GB GPU device memory

4 x Xilinx Virtex-6 FPGA (ML 605)

since beg. 2013 operated.



Nr. 1 Supercomputer from China: 33 Pflop/s Linpack

Tianhe-2 (MilkyWay-2) - TH-IVB-FEP Cluster, Intel Xeon E5-2692 12C 2.200GHz, TH Express-2, Intel Xeon Phi 31S1P

<http://www.top500.org>



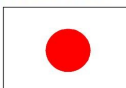



32000 Intel Xeon 12 core, 48000 Intel Phi Accelerators 57 Core

Top 10 List November 2010

From www.top500.org - list of fastest

supercomputers in the world...
... last year Nov. 2010:

1	National Supercomputing Center in Tianjin China		Tianhe-1A - NUDT TH MPP, X5670 2.93Ghz 6C, NVIDIA GPU, FT-1000 8C NUDT	<u>GPU</u>
2	DOE/SC/Oak Ridge National Laboratory United States		Jaguar - Cray XT5-HE Opteron 6-core 2.6 GHz Cray Inc.	
3	National Supercomputing Centre in Shenzhen (NSCS) China		Nebulae - Dawning TC3600 Blade, Intel X5650, NVidia Tesla C2050 GPU Dawning	<u>GPU</u>
4	GSIC Center, Tokyo Institute of Technology Japan		TSUBAME 2.0 - HP ProLiant SL390s G7 Xeon 6C X5670, Nvidia GPU, Linux/Windows NEC/HP	<u>GPU</u>
5	DOE/SC/LBNL/NERSC United States		Hopper - Cray XE6 12-core 2.1 GHz Cray Inc.	
6	Commissariat a l'Energie Atomique (CEA) France	FR	Tera-100 - Bull bullx super-node S6010/S6030 Bull SA	
7	DOE/NNSA/LANL United States		Roadrunner - BladeCenter QS22/LS21 Cluster, PowerXCell 8i 3.2 Ghz / Opteron DC 1.8 GHz, Voltaire Infiniband IBM	
8	National Institute for Computational Sciences/University of Tennessee United States		Kraken XT5 - Cray XT5-HE Opteron 6-core 2.6 GHz Cray Inc.	
9	Forschungszentrum Juelich (FZJ) Germany		JUGENE - Blue Gene/P Solution IBM	
10	DOE/NNSA/LANL/SNL United States		Cielo - Cray XE6 8-core 2.4 GHz Cray Inc.	

► China Grabs Supercomputing Leadership Spot in Latest Ranking of World's Top 500 Supercomputers

Thu, 2010-11-11 22:42

MANNHEIM, Germany; BERKELEY, Calif.; and KNOXVILLE, Tenn.—The 36th edition of the closely watched TOP500 list of the world's most powerful supercomputers confirms the rumored takeover of the top spot by the Chinese Tianhe-1A system at the National Supercomputer Center in Tianjin, achieving a performance level of 2.57 petaflop/s (quadrillions of calculations per second).

Ranking the World's Most ENERGY-EFFICIENT SUPERCOMPUTERS

www.green500.org



Green500 Rank	MFLOPS/W	Site*	Computer*	Total Power (kW)
1	4,503.17	GSIC Center, Tokyo Institute of Technology	TSUBAME-KFC - LX 1U-4GPU/104Re-1G Cluster, Intel Xeon E5-2620v2 6C 2.100GHz, Infiniband FDR, NVIDIA K20x	27.78
2	3,631.86	Cambridge University	Wilkes - Dell T620 Cluster, Intel Xeon E5-2630v2 6C 2.600GHz, Infiniband FDR, NVIDIA K20	52.62
3	3,517.84	Center for Computational Sciences, University of Tsukuba	HA-PACS TCA - Cray 3623G4-SM Cluster, Intel Xeon E5-2680v2 10C 2.800GHz, Infiniband QDR, NVIDIA K20x	78.77
4	3,185.91	Swiss National Supercomputing Centre (CSCS)	Piz Daint - Cray XC30, Xeon E5-2670 8C 2.600GHz, Aries interconnect, NVIDIA K20x Level 3 measurement data available	1,753.66
5	3,130.95	ROMEO HPC Center - Champagne-Ardenne	romeo - Bull R421-E3 Cluster, Intel Xeon E5-2650v2 8C 2.600GHz, Infiniband FDR, NVIDIA K20x	81.41
6	3,068.71	GSIC Center, Tokyo Institute of Technology	TSUBAME 2.5 - Cluster Platform SL390s G7, Xeon X5670 6C 2.930GHz, Infiniband QDR, NVIDIA K20x	922.54
7	2,702.16	University of Arizona	iDataPlex DX360M4, Intel Xeon E5-2650v2 8C 2.600GHz, Infiniband FDR14, NVIDIA K20x	53.62
8	2,629.10	Max-Planck-Gesellschaft MPI/IPP	iDataPlex DX360M4, Intel Xeon E5-2680v2 10C 2.800GHz, Infiniband, NVIDIA K20x	269.94
9	2,629.10	Financial Institution	iDataPlex DX360M4, Intel Xeon E5-2680v2 10C 2.800GHz, Infiniband, NVIDIA K20x	55.62
10	2,358.69	CSIRO	CSIRO GPU Cluster - Nitro G16 3GPU, Xeon E5-2650 8C 2.000GHz, Infiniband FDR, Nvidia K20m	71.01

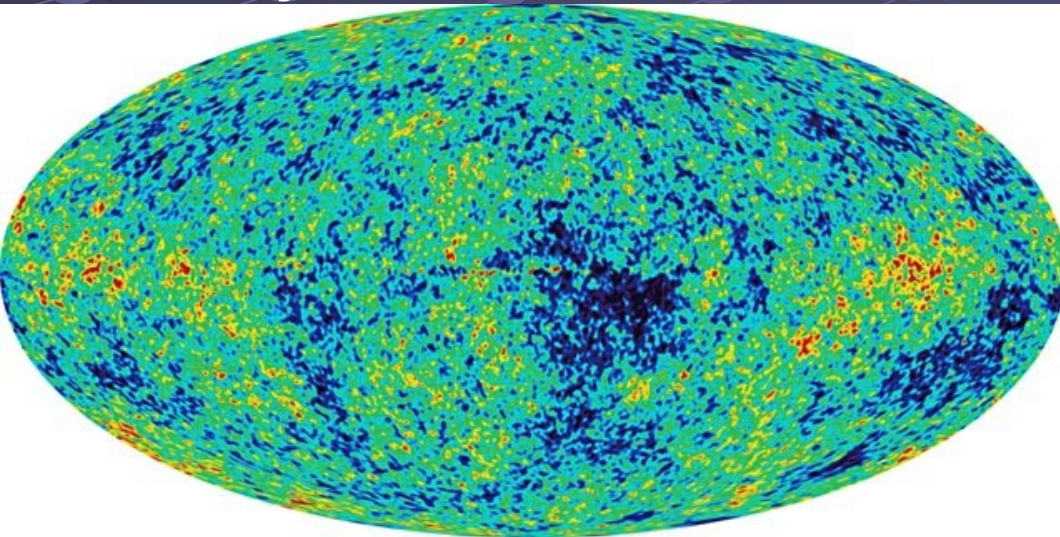
Computer Physics - Astrophysics

Cosmology

Computer Physics – Astrophysics

● Structurr Formation in the Universe

In the year 100.000....



● Wilkinson Microwave Anisotropy Probe (WMAP)
(Cosmic Microwave Background)

...and ``today``

A visualization of the Millennium Simulation, showing a vast field of particles. The particles are represented as small dots, with colors ranging from blue to red, indicating different temperatures or velocities. The overall structure is a complex, interconnected network of filaments and voids, characteristic of the cosmic web. A horizontal double-headed arrow is positioned above the text '1 Gpc/h', indicating the scale of the simulation.

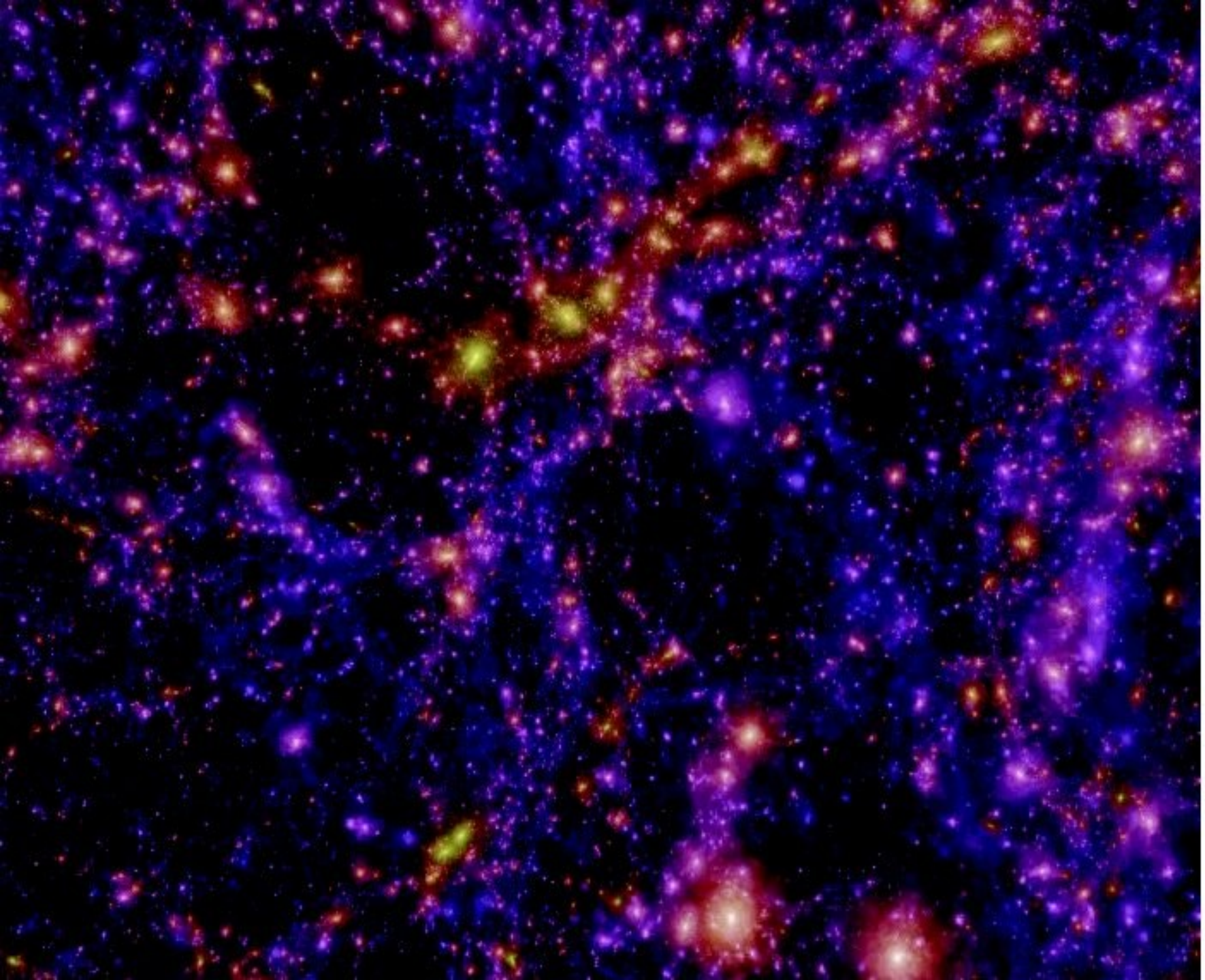
1 Gpc/h

Millennium Simulation

10,077,696,000 particles

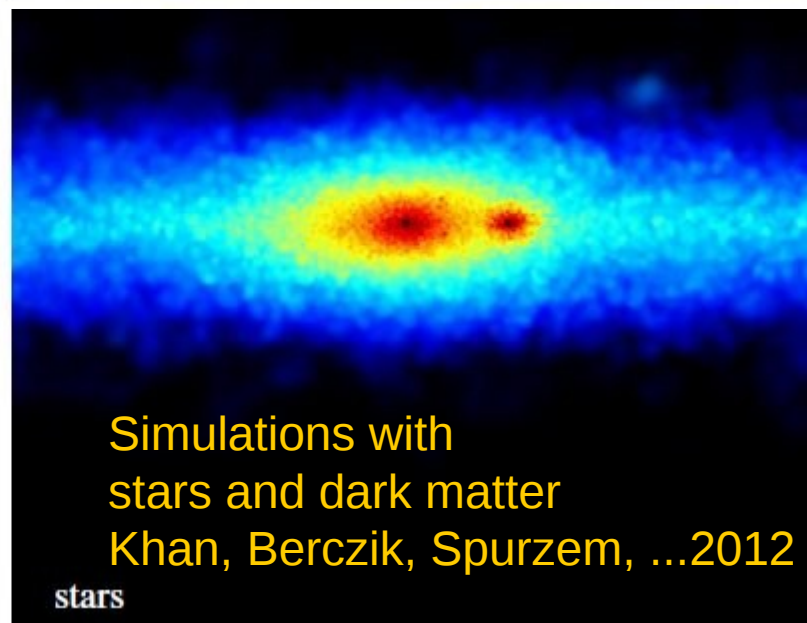
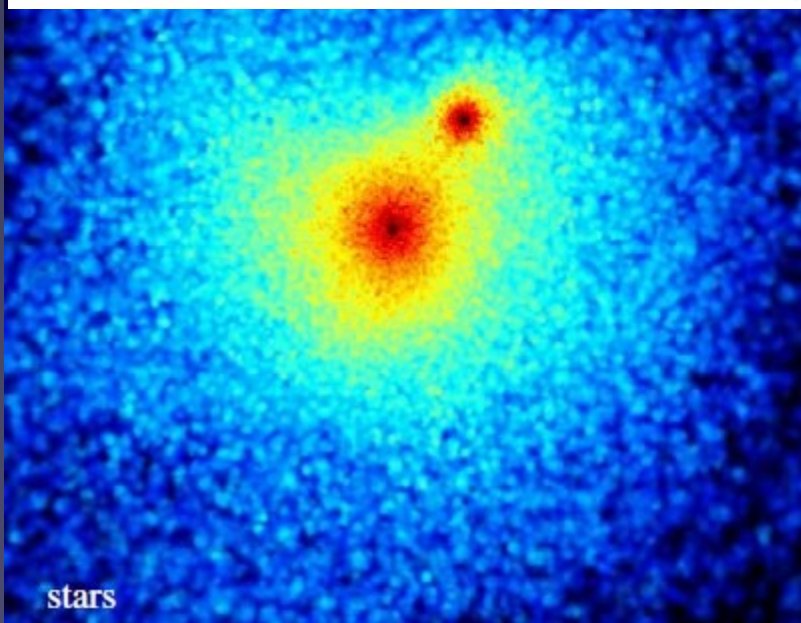
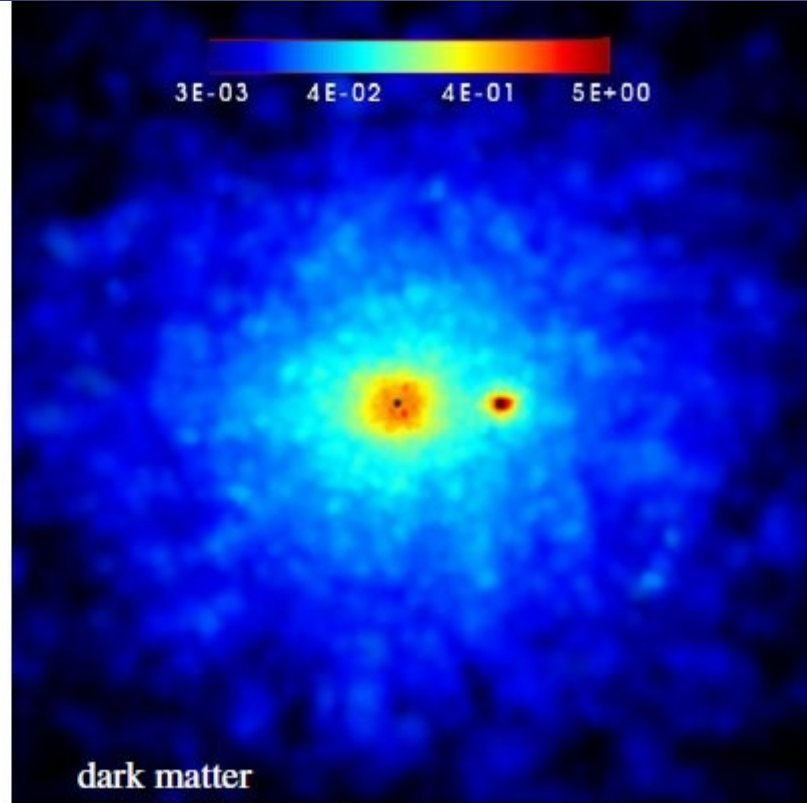
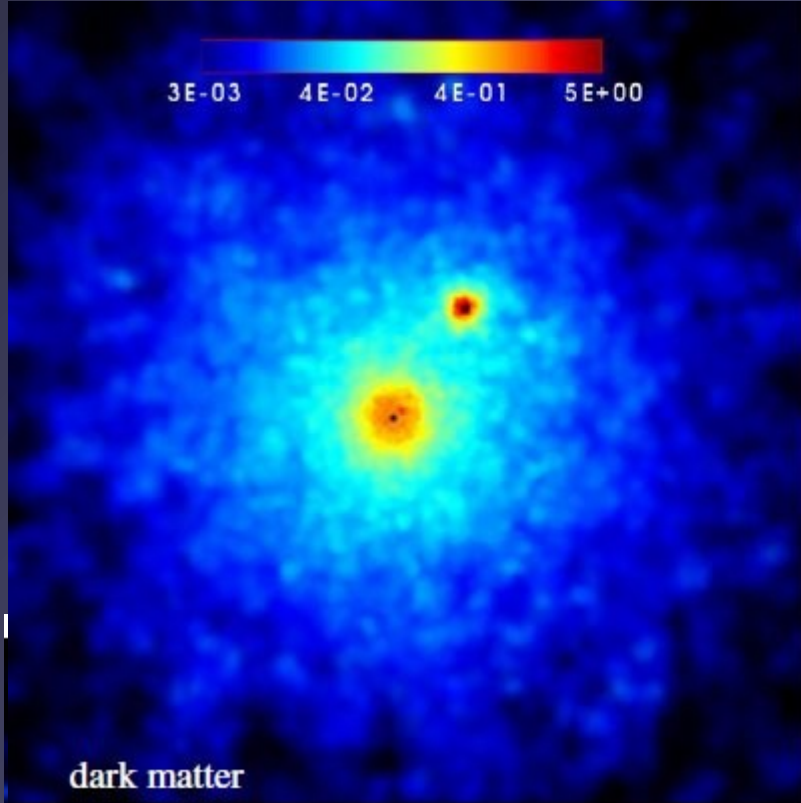
($z = 0$)

Millennium Simulation (Springel et al.)

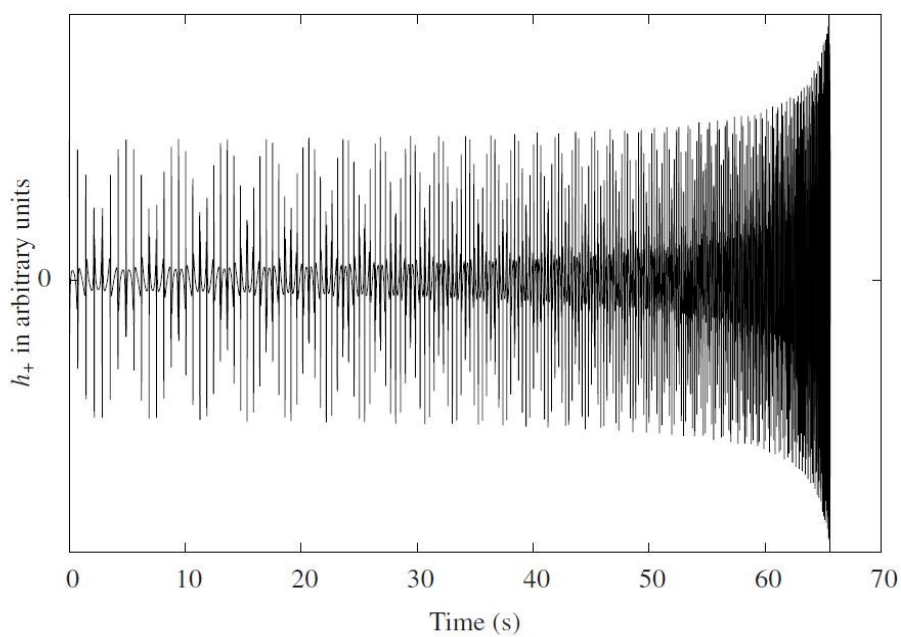


Computer Physics - Astrophysics

Black Holes in Galaxies



Simulations with stars and dark matter
Khan, Berczik, Spurzem, ...2012



Post-Newtonian Dynamics Gravitational Wave Templates

Figure 3.11: Waveform for two equal mass objects on a an orbit with $e = 0.5$.

Handle spin-orbit and
spin-spin coupling
(P.Brem, R. Spurzem,
Univ. Heidelberg)

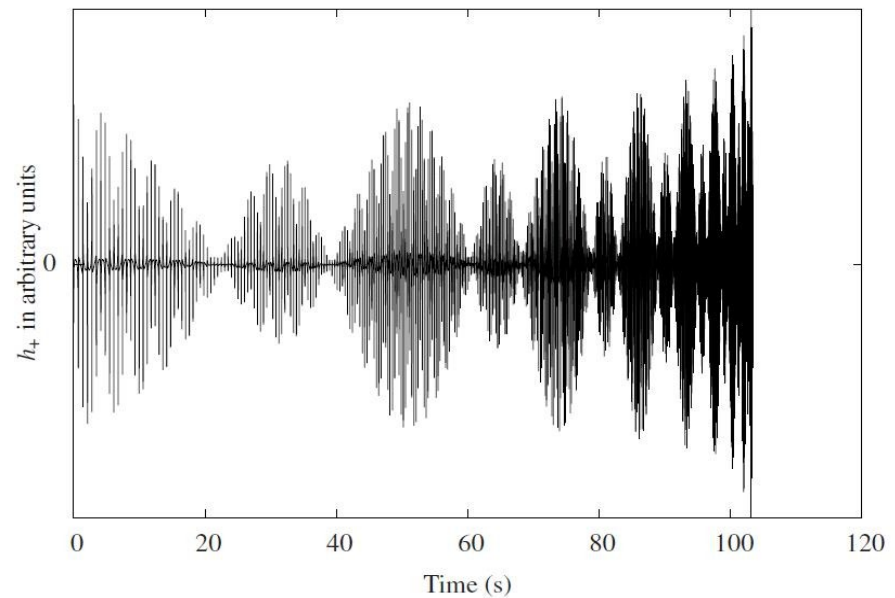


Figure 3.12: Waveform for two objects with a mass ratio of $q = 1/10$ on an orbit with $e = 0.5$ and spins $a_{1,x} = 1.0$, $a_{2,y} = 1.0$.

EUROPEAN GRAVITATIONAL OBSERVATORY

EGO



Consortium of

Example: VIRGO Detector in Cascina near Pisa, Italy





VIRGO – Pisa 3km
LIGO – Livingston, LA
Hanford, WA
1km
GEO600 – Hannover
600m
AIGO – Australien
(planned, 5 km)

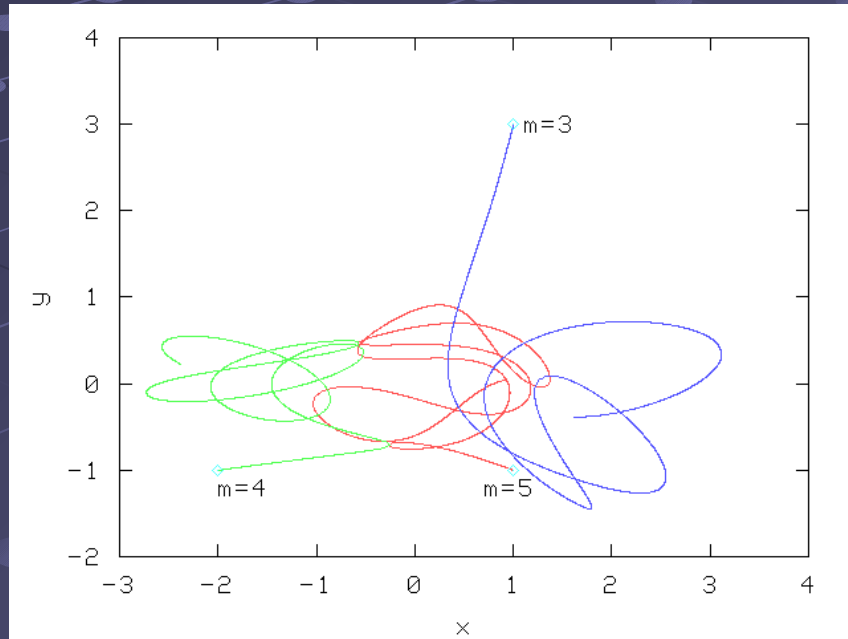
<http://www.ligo-la.caltech.edu/>
<http://www.ego-gw.it>
<http://www.geo600.uni-hannover.de>

Outreach to 50 Millionen
light years (Neutron Stars)

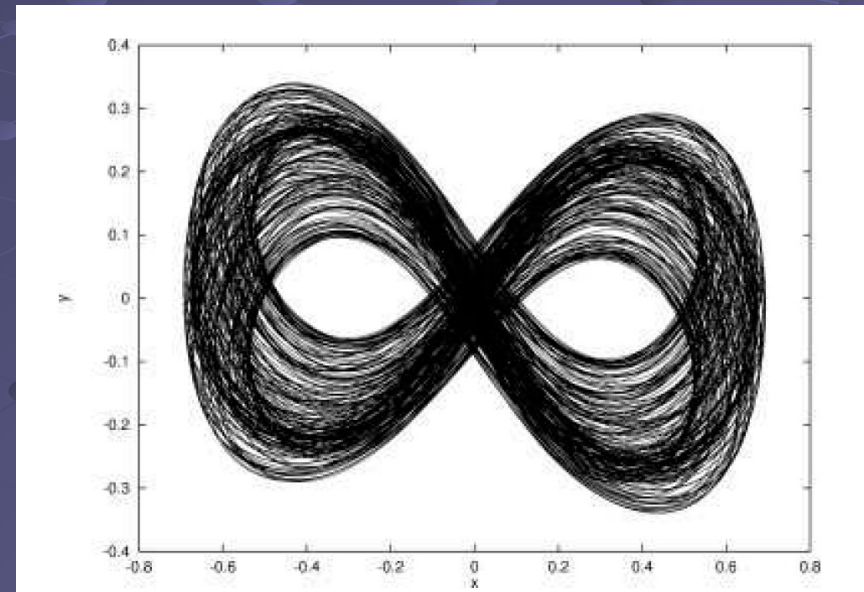
Nonlinear Dynamics and Chaos

3-Body Problems

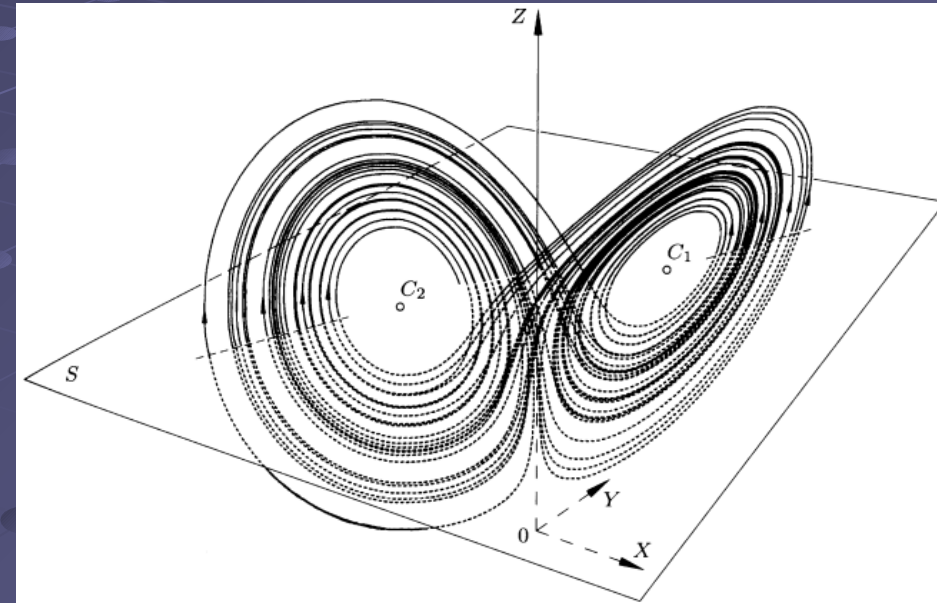
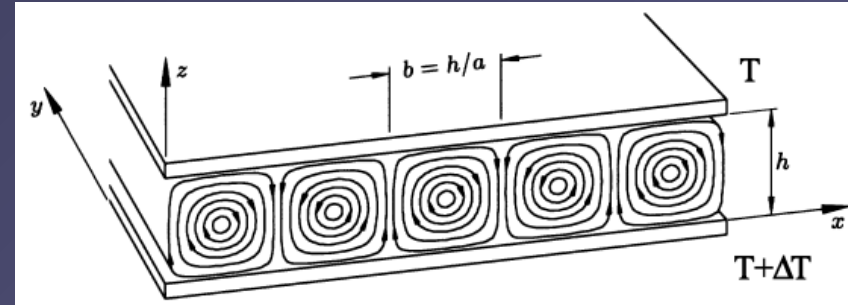
Burrau's Problem



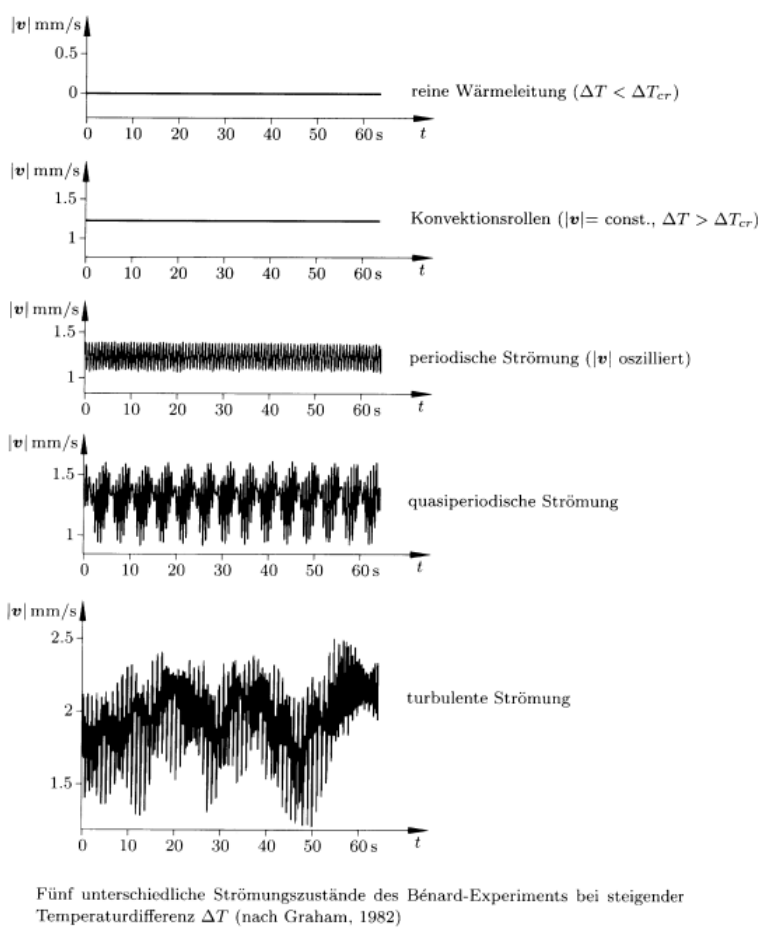
„The Eight“



Lorenz-Attraktor



Lorenz-Attraktor für $r = 28$, $\sigma = 10$ und $b = 8/3$.
 Der Trajektorienbereich, den die Ebene $Z = r - 1 = 27$ verdeckt, ist punktiert (Lanford, 1977)



Logistische Abbildung

